

RESULTS

OF THE

MAGNETICAL AND METEOROLOGICAL

OBSERVATIONS

HADE AT

THE ROYAL OBSERVATORY, GREENWICH,

1870.

(EXTRACTED FROM THE GREENWICH OBSERVATIONS, 1870.)

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GREENWICH MAGNETICAL AND METEOROLOGICAL OBSERVATIONS, 1870.

INTRODUCTION.

§ 1. *Buildings qjthe Magnetic Observatory.*

IN consequence of a representation by the Astronomer Roya1, dated 1836, January 12, and a memorial by the Board of Visitors of the Royal Observatory, dated 1836, February 26, addressed to the Lords Commissioners of the Admiralty, an additional space of ground on the south-east side of the former boundary of the Observatory grounds was inclosed from Greenwich Park for the site of a Magnetic Observatory, in the summer of 1837, and the Magnetic Observatory was erected in the spring of 1838. Its nearest angle in its present form is about 174 feet from the nearest point of the S.E. dome, and about 30 feet from the office of Clerk of Works. It is based on concrete and built of wood, united for the most part by pegs of bamboo; no iron was intentionally admitted in its construction, or in subsequent alterations. Its form, as originally built, was that of a cross with four equal arms, very nearly in the direction of the cardinal magnetic points as they were in 1838; the length within the walls, from the extremity of one arm of the cross to the extremity of the opposite arm, was 40 feet, the breadth of each arm 12 feet. In the spring of 1862, the northern arm was extended 8 feet. The height of the walls inside is 10 feet, and the ceiling of the room is about 2 feet higher. The northern arm of the cross is separated from the central square by a partition, so as to form an ante-room. The meridional magnet, for observations of absolute declination and of variations of declination (placed in its position in 1838), is mounted in the southern arm; and the theodolite by which the magnet collimator is viewed, and by which circumpolar stars for determination of the astronomical meridian are also observed (for which observation an opening is made in the roof, with proper shutters,) is in the southern arm, near the southern boundary of the central square. The bifilar magnet, for variations of horizontal magnetic force (erected at the end of 1840) was mounted near the northern wall of the eastern arm; and the balance-magnetometer, for variations of vertical magnetic force (erected in 1841) was mounted near the northern wall of the western arm. Important changes have subsequently been made in the positions of these instruments, as will be mentioned below. The sidereal time-clock is in the south arm, near the southeast re-entering angle. The fire-grate (constructed of copper, as far as possible,) is near the north end of the west side of the ante-room. Some of these fixtures may contain trifling quantities of iron, and, as the ante-room is used as a computing room

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it is impossible to avoid the introduction of iron in small quantities; great care, however, is taken to avoid it as far as possible.

In 1864, a room, called the Magnetic Basement, was excavated below the whole of the Magnetic Observatory except the ante-room; the descent to it is by a staircase close to the south wall of the western arm of the building.

For the theodolite, a brick pier was built from the ground below the floor of the Basement, rising through the ceiling into the south arm of the upper room, and supporting the theodolite in exactly the same position as before.

Instead of a single meridional magnet performing the double functions of "magnet for determining absolute magnetic declination," and "magnet carrying a rnirror for photographic register," there are now two meridional magnets, one in the Upper Room and one in the Basement. The upper magnet is in a position about 10 inches north of the former position of the declination-magnet; it carries a collimator, for observation by the theodolite; but, in reversion of position of the collimator, the collimator is always either above or below the magnet, so that the magnet is always in the same vertical. The lower magnet, which is in the same vertical with the upper magnet, carries the mirror for the photographic register of the continual changes of declination. A massive brick pier is built in the south arm of the Basement, covered by a stone slab; upon it is fixed the . gun-metal stand carrying the photographic lamp, and the narrow chink through which it shines; from the stone slab rise three smaller piers, upon which crossed slates are placed; and from these rises a small pier through the ceiling, to the height of 18 inches above the upper floor, carrying the suspension of the lower magnet; the skein of silk, which supports the lower magnet, passes through a hole in one of the slates. Upon the tops of the three piers rest the feet of the original wooden stand carrying the suspension of the upper magnet.

The bifilar-magnetometer is in the Basement, in a position vertically below its former position. A massive brick pier, surmounted by a thick slab of stone (upon which the metal stand carrying the photograph lamp and narrow chink is fixed) carries a pier consisting of a back and return-sides, which rises through the ceiling about 2 feet above the upper floor, and is crowned by a slate slab that carries the suspension of the bifilar-magnetometer.

The vertical-force magnetometer is in the Basement, in a position vertically below its former position; it rests upon a brick pier, capped by a thick stone; to which also is fixed the plate of metal with narrow chink through which passes the light of the photographic lamp.

To the theodolite-pier are fixed telescopes for eye-observation of the bifilar and vertical-force magnetometers.

At the south-east re-entering angle of the Basement (which has been rebated for the purpose) is the horizontal photographic cylinder, which receives the traces of the movements of the declination-magnet and the bifilar-magnet. The angle is so far cut away that the straight line joining their suspensions passes at the distance of one foot from the wall, and thus the cylinder receives the light from the concave mirrors carried by both instruments, at right angles to its surface. The vertical cylinder

BUILDINGS OF THE MAGNETIC OBSERVATORY.

which receives the traces of the movements of the vertical-foree-magnet, and of the self-registering barometer near it, is east of the vertical force pier.

In the south-west corner of the western arm, and partially beneath the staircase, was placed, until 1870, May 8, the apparatus for self-registration of the spontaneous galvanic currents on the wires leading respectively, from Angerstein Wharf to Lady Well Station (on the Mid Kent Railway), and from North Kent Junction (on the Greenwich Railway) to Morden College end of the Blackheath Tunnel (on the North Kent Railway). The straight lines connecting these points intersect each other nearly at right angles, at a point not far distant from the Observatory (see \S 13 below). The self-registering apparatus is now transferred to the south-eastern corner of the eastern arm.

The mean-time-clock is on the west wall of the south arm of the Basement.

Adjoining the north wall is the table for photographic operations. \cdot Much water is used in these operations, and therefore a pump is provjded in the grounds at a distance of about 30 feet from the nearest magnetometer, by which the water is withdrawn from the cistern at the east end of the photographic table and at once discharged into a covered drain.

The Basement is warmed by a gas-stove, and ventilated by a large copper tube nearly two feet in diameter, receiving the flues from the stove and all the lamps, and passing through the upper room to a revolving cowl above the roof. Each of the arms of the basement has a window facing the south, but in general the window-wells are closely stopped.

The variations in the temperature of the instruments have been greatly reduced by their location within this Basement.

On the outside of the Magnetic Observatory, near the north-east corner of the ante-room, a pole 79 feet in height is fixed, for the support of the conducting wires to the electrometers; the electrometers, &c., are planted in the window-seat at the north-end of the ante-room.

The apparatus for naphthalizing the gas used in the photographic registration was formerly fixed in a corner of the ante-room, but is now (1870) mounted in a small detached zinc-built room, erected in 1863, near the west side of the ante-room. No naphthalizing process, however, has been in use since the year 1865.

In 1863, a range of seven rooms, usually called the Magnetic Offices, was erected near the southern fence of the grounds. Since the summer of 1863, observations of Dip and Deflexion have been made in the westernmost of these rooms.

At the distance of 28 feet south (magnetic) from the south-east angle of the southern arm is a square shed about $10th 6ⁱⁿ$ square, supported by four posts at the height 8 feet, with an adjustible opening at the center of the top. Under this shed are placed the large dry-bulb and wet-bulb thermometers, with a photographic cylinder, whose axis is vertical, between them; and external to these are the gas flames, whose light passing through the thermometer-tubes above the quicksilver makes photographic traces upon the paper which covers the cylinder.

For better understanding of these descriptions, the reader is referred to the Descriptions of Buildings and Grounds with accompanying Maps, attached to the Volumes of Astronomical Observations for the years 1845 and 1862.

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§ 2. *Upper Declination-Magnet and Apparatus for observing it.*

The theodolite with which the meridional magnet is observed is by Simms: the radius of its horizontal circle is 8.3 inches: it is divided to 5', and reads to 5'', by three verniers, carried by the revolving frame of the theodolite. The fixed frame stands upon three foot-screws, which rest in brass channels let into a stone pier, that stands upon the brick pier rising from the ground of the Magnetic Basement. The revolving frame carries the Y's (with vertical adjustment at one end) for a telescope with transit-axis: the length of the axis is $10\frac{1}{2}$ inches: the length of the telescope 21 inches: the aperture of the object glass 2 inches. The Y's are not carried immediately by the T head which crosses the vertical axis of the revolying frame, but by pieces supported by the ends of that T head, and projecting horizontally from it: the use of this construction is to allow the telescope to be pointed sufficiently high to see δ Ursæ Minoris above the pole. The eye-piece of the telescope carries only one fixed horizontal wire, and one vertical wire moved by a micrometer-screw. The opening in the roof of the building permits the observation of circumpolar stars, as high as δ Ursæ Minoris above the pole, and as low as β Cephei below the pole.

For supporting the magnet, a braced wooden tripod-stand is provided, whose feet, as above described, rest upon brick piers in the Magnetic Basement. Upon the cross-bars of the stand rests a double rectangular box (one box completely inclosed within another), both boxes being covered with gilt paper on their exterior and interior sides. On the southern side of the principal upright piece of the stand is a moveable upright bar, turning in the vertical E. and W. plane, upon a pin in its center (which is fixed in the principal upright), and carrying at its top the pulleys for suspension of the magnet; this construction is adopted as convenient for giving an E. and W. movement (now very rarely required) to the point of suspension, by giving a motion to the lower end of the bar. The top of the upright piece carries a brass frame with two pulleys, whose axes are E. and W.: one of these pulleys projects beyond the north side of the principal upright, and from it depends the suspension skein: the other pulley projects on the south side: the suspension skein, being brought from the magnet up to the north pulley, is carried over it and over the south pulley, and thence downwards to a small windlass, carried by the lower part of the moveable upright. The height of the two pulleys above the floor is about 11 ft. $3\frac{3}{4}$ in., and the height of the magnet is about 2 ft. 10 in.; the length of the metal carrier which bears the magnet is 1 ft. 3 in.; so that the length of the free suspending skein is about 7 ft. $2\frac{3}{4}$ in.

The magnet was made by Meyerstein, of Göttingen: it is a bar 2 feet long, $1\frac{1}{2}$ inch broad, and about $\frac{1}{6}$ inch thick: it is of hard steel throughout. The magnet-carrier was also made by Meyerstein, but it has since been altered by Simms. The magnet is inserted sideways and fixed by screws in a double square hook which constitutes the lower part of the magnet-carrier. This lower part turns stifHy by a vertical axis with index in a graduated horizontal circle (usually called the torsion-circle) attached to the upper part. The upper part of the magnet-carrier is simply hooked into the skein.

The suspending skein was originally of silk fibre, in the state in which it is first

prepared by silk manufacturers for further operations; namely, when seven or more fibres from the cocoon are united by juxtaposition only (without twist) to form a single thread. The skein was strong enough to support perhaps three times the weight of the magnet, &c.

In the summer and autumn of 1864, an attempt was made to suspend the magnet by a steel wire, capable of supporting the weight 15 lbs.; but the torsion force was found to be so large as greatly to diminish the value of the observations; and the skein was finally restored on 1865, January 20. A similar attempt was made for suspension of the lower magnet; the skein, however, was restored on 1865, January 30.

Upon the magnet there slide two brass frames, firmly fixed in their places by means of pinching-screws. One of these contains, between two plane glasses, a cross of delicate cobwebs; the other holds a lens of 13 inches focal length and nearly 2 inches aperture. This combination, therefore, serves as a reversed telescope without a tube: the cross of cobwebs is seen very well with the theodolite-telescope, when the suspensionbar of the magnet is so adjusted as to place the object-glass of the reversed telescope in front of the object-glass of the theodolite, their axes coinciding. The wires are illuminated by a lamp and lens in the night, and by a reflector in the day.

In the original mounting of this magnet the small vibrations were annihilated by a copper oval or "damper," thus constructed: A copper bar, about one inch square, is bent into a long oval form, intended to contain within itself the magnet (the plane of the oval curve being vertical). A lateral bend is made in the upper half of the oval, to avoid interference with the suspension-piece of the magnet. The effect of this damper was, that after every complete or double vibration of the magnet, the amplitude of the oscillation is reduced in the proportion of 5 : 2 nearly.

On mounting the photographic magnetometer in the basement, the damper was removed from its place surrounding the upper magnet, and was adjusted to encircle the photographic magnet. The upper magnet remained unchecked in its vibrations till 1866, January 23, when the lower part of its magnet-carrier was connected with a brass bar which vibrates in water.

OBSERVATIONS RELATING TO THE PERMANENT ADJUSTMENTS OF THE UPPER DECLINATION-MAGNET AND ITS THEODOLITE.

1. Determination of the inequality of the pivots of the theodolite-telescope.

1862, December 26. The theodolite was clamped, so that the transit-axis was at right angles to the astronomical meridian. The illuminated end of the axis of the telescope was first placed to the East: the level was applied, and its scale was read; the level was then reversed, and its scale was again read; it was then again reversed, and again read, and so on successively six times. The illuminated end of the axis was then placed to the West, and the level was applied and read as before. This process was repeated four times, and the result was that, when the level indicates the axis to be horizontal, the pivot at the illuminated end is really too low by 0"'3 nearly.

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2. Value of one revolution of the micrometer-screw of the theodolite telescope.

On 1862, December 26, observations were made, giving for the value of one revolution of the micrometer $1'.33''·85.$ On 1865, December 27, the magnet was made to rest on blocks of wood, and its collimator was used as a fixed mark at an infinite distance. The micrometer of the theodolite was placed in different positions, and the telescope of the theodolite was then turned till the micrometer wire bisected the cross. The result of ten comparisons of theodolite-readings with large values and with small values of the micrometer-reading was, that one revolution $= 1'$. $34''·8$. This is used through the year 1870.

3. Determination of the micrometer-reading for the line of collimation of the theodolite-telescope.

1869, December 28. The vertical axis of the theodolite had been adjusted to verticality, and the transit-axis was made horizontal. The declination-magnet was made to rest on blocks, and the cross-wires carried by it were used as a collimator for determining the line of collimation of the telescope of the theodolite. The telescope' was reversed after each observation. The mean of 20 double observations was 100^r·084. This value is used throughout the year 1870.

4. Determination of the effect of the mean-time-clock on the declination-magnet.

The observations by which this has been determined are detailed in the volumes for 1840, 1841, 1844, and 1845. It appeared that it was necessary to add 9"-41 to every reading of the theodolite. The clock was removed to the basement in 1864, having now nearly the same relative position to the lower declination-magnet which formerly it had to the upper. No correction is now applied to the upper deciination-magnet.

5. Determination of the compound effects of the vertical-force-magnet and the horizontal-force-magnet on the declination-magnet.

The details applying to the effect of the horizontal-foree-magnet and first' verticalforce-magnet will be found in the volumes for 1840 , 1841 , 1844 , and 1845 . It appeared that it was necessary to subtract $55^{\prime\prime}$: 22 from all readings of the theodolite. In 1848 a new vertical-foree-magnet was introduced, and the subtractive quantity was then found to be 42"·2. A few experiments in 1865 seemed to show that the correction is now 36".9. No numerical correction has been applied.

6. Determination of the error of collimation for the plane glass in front of the boxes of the declination-magnet.

1869, December 28. The magnet was made to rest entirely on blocks. The micrometer head of the telescope was to the East. The plane glass has the word " top" engraved on it, and, in ordinary use, this word is always kept east. The cross-wire carried by the collimator of the magnet was observed with the engraved word alternately east and west. The result of 20 double observations was, that in the ordinary position of the glass 19° ; is to be added to all readings.

7. Determination of the error of collimation of the magnet-collimator, with reference to the magnetic axis of the magnet.

1869, December 28. Observations were made by placing the declination-magnet

in its stirrup, with its collimator alternately above and below, and observing the collimator-wire by the theodolite-telescope; the windlass of the suspending skein being so moved that the collimator in each observation was in the line of the theodolitetelescope. Seven pairs of observations were taken. The mean half excess of reading with collimator above, (its usual position) over that with collimator below was $25'$. 54" $4'$. The value used in the reductions for 1870 is $25'$. $25''$ 7 (the mean of the results for the five years 1866-1870).

8. Effect of the damper.

In the volume for 1841 observations are exhibited shewing that the oval copper bar, or damper, which then surrounded what is now the upper declination-magnet, had but little 'or no effect. Repeated observations, of less formal character, in succeeding years, have confirmed this result. The same bar has encircled the lower declinationmagnet since the year 1865. The following observations were made in the year 1865, for ascertaining the effect of the damper on the lower declination-magnet under various circumstances.

On 1865, February 8 and 10, and March 2, the time of vibration of the magnet was $observed :=$

These seem to indicate a repulsion of the magnet by the damper, but the magnet came to rest so rapidly that the observations are very uncertain:

On several days from 1865, April 2 to May 12, observations were made for ascertaining the deflexion of the magnet produced by turning the damper through a small angle round a vertical axis, passing through its center.

The first series shews clearly that the damper in its usual position drags the magnet; the second shews no certain effect. It seems that the damper possesses two kinds of GREENWICH MAGNETICAL AND METEOROLOGICAL OBSERVATIONS, 1870. *h*

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magnetism, one permanent, the other transiently induced, of nearly equal magnitude; their sum being about $\frac{1}{100}$ part of the terrestrial effect for the same deflexion.

From 1865, July 25 to August 9, observations were made to ascertain whether the effect of an external deflecting cause is the same with the damper present and the damper removed. The observation was extremely difficult, as the magnet was perpetually in vibration when the damper was removed. A small magnet on the east side of the N. end of the magnetometer, with its north end pointing towards the East (and therefore diminishing the western declination of the magnetometer), was moved to the distance (about five feet) at which it produced a deviation of 5' nearly. The apparent western declination was observed, damper present, and damper removed. It appeared to be less with damper present than with damper removed, by 0'. 53". The separate results are very discordant. If the conclusion has any validity, it tends to shew a repulsive power in the damper, opposite to that found in the preceding experiments. This experiment is regarded as inconclusive.

9. Calculation of the constant used in the reduction of the observations of the upper declination-magnet, the micrometer-head of the theodolite-telescope being East.

Constant to be used in the reduction of the observations $\dots \dots \dots \dots \dots -3.$ 3. 14.0

10. Determination of the time of vibration of the upper declination-magnet under the action of terrestrial magnetism.

On 1868, January 22, it was found to be $30^{\circ}60$; on March 19, $30^{\circ}56$; on December 30, 30°50; and on 1869, November 13, 30°50.

11. Fraction expressing the proportion of the torsion-force to the earth's magnetic force.

By the same process which is described in the Magnetical Observations 1847, but with the silk skein now in use, the proportion was found, on 1865, January 31, $\frac{1}{214}$; on February 17, $\frac{1}{227}$; on April 27, $\frac{1}{207}$; on December 27, $\frac{1}{230}$; and on 1869, December 29, $\frac{1}{262}$.

DETERMINA TION OF THE READINGS OF THE HORIZONTAL CIRCLE OF THE THEODOLITE CORRESPONDING TO THE ASTRONOMICAL MERIDIAN.

The error of the level is determined by application of the spirit-level at the time of observation: due regard being paid, in the reduction, to the inequality of pivots already found. One division of the level is considered $= 1"0526$. The azimuthreading is then corrected by this quantity;

Correction $=$ Elevation of W. end of axis \times tan star's altitude.

The readings of the azimuth circle increase as the instrument is turned' from N. to E., S., and W.; from which it follows that the. correction must have the same sign as the elevation of the W. end.

The correction for the azimuth of the star observed has been computed independently in every observation, by a peculiar method, of which the principle is fully explained in the volumes for 1840-1841, 1843, 1844, 1845. The formula and table used are the following $:$ ---

Let A_{μ} = seconds of arc in star's azimuth,

 C_s = seconds of time in star's hour-angle,

 a_n = seconds of arc in star's N.P.D. for the day of observation,

Then log. $A_{\nu} = \log C_s + \log E + \log (a_{\nu} + F) + \log \cos \phi$.

The values of log. E, F, and log. cos φ , are given in the following table :-

TABULATED VALUES of LOG. Cos ϕ , for DIFFERENT VALUES of C_s , and of the QUANTITIES Log. E and F, for the STARS POLARIS and δ URSE MINORIS.

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Observations for determining the theodolite readings corresponding to the astronomical meridian were made on the following days in 1870: -- January 14, 26; February 10, 24; March 5; April 1, 5, 14, 18; May 9, 12, 14, 16, 18, 19, 21, 25; June 1, 14, 21; July 12,15, 23; August II, 29,31; October 10, 11, 14; November 12, 17; December 14, 24. As a check on the continued steadiness of the theodolite, observations of a fixed mark (a small hole in a plate of metal above the Observatory Library, illuminated by a reflector of sky-light in the day and by a lamp at night,) have been taken about twenty times at nearly equal intervals through the year.

The following is a description of the method of making and reducing the eye observations of the declination-magnet $:$

A fine horizontal wire (as stated above) is fixed in the field of view of the theodolitetelescope, and another fine vertical wire is fixed to a wire-plate, moved right and left by a micrometer screw. On looking into the telescope, the cross of the magnetometer is seen; and during the vibration of the magnet, this cross is seen to pass alternately right and left. The observation is made by turning the micrometer till its wire bisects the image of the magnet-cross at the pre-arranged times, and reading the micrometer. The verniers of the horizontal circle are read.

The mean-time clock is kept very nearly to Greenwich mean time (its error being ascertained each day), and the clock-time for each determination is arranged beforehand. Chronometer M^cCabe 649 has usually been employed for observation.

If the magnet is in a state of disturbance, the first observation is made by the observer applying his eye to the telescope about one minute before the pre-arranged time; he bisects the magnet-cross by the micrometer wire at 45^s, and again at 15^s before that time, also at 15^s and 45^s after that time. The intervals of these four observations are therefore the same as the time of vibration of the magnet, and the mean of all the times is the same as the Greenwich pre-arranged mean time.

The mean of each pair of adjacent readings of the microineter is taken (giving three means), and the mean of these three is adopted as the result. In practice, this is done by adding the first and fourth readings to the double of the second and third, and dividing the sum by 6.

Till 1866, January 23, the magnet was usually in a state of vibration; but, since the introduction of the water-damper on that day, the number of instances of vibration has been very small. When it is found to be quite free from vibration, two bisections only of the cross are made, one about 15^s before the time recorded, the other about 15^s after that time, 30^s being nearly the time of a single vibration. (The lower magnet, furnished with the copper damper, never exhibits any troublesome vibrations.)

The adopted result is converted into arc, supposing $I^r = I'$. 34".8, and the quantity thus deduced is added to the mean of the vernier-readings, from which is subtracted the constant given in article 9 of the permanent adjustments; the difference between this number and the adopted reading for the Astronomical South Meridian is taken;

EYE-OBSERVATIONS OF DECLINATION MAGNET. GENERAL PRINCIPLE OF PHOTOGRAPHIC REGISTRATION.

and thus is deduced the magnetic declination, which is used in determining the zero for the photographic register.

§ 3. *General principle* of *construction* of *Photographic self-registering Apparatus for continuous Record* of *Magnetic and other Indications.*

The general principle adopted for all the photographic instruments is the same. For the register of each indication, a cylinder is provided, whose material is ebonite, and which is very accurately turned in the lathe. The axis of the cylinder is placed . parallel to the direction of the change of indication which is to be registered. If there are two indications whose movements are in the same direction, both may be registered on the same cylinder; thus, the Declination and the Horizontal Force, whose indications of changes of the respective elements are both made to travel horizontally, can both be registered upon one cylinder with axis horizontal: the same remark applies to the register of two different galvanic Earth-Currents; the Vertical Force and the reading of the Barometer can both be registered upon one cylinder with axis vertical; and similarly the Dry-Bulb Thermometer and the Wet-Bulb Thermometer.

To the ends of each ebonite cylinder there are fixed circular brass plates, that which is near the clock-work having a diameter somewhat greater than that of the cylinder. In the further fittings there is a little difference between those for vertical and those for horizontal cylinders. Each horizontal cylinder has a pivot fixed in the brass plate at each end; these revolve each upon two antifriction wheels of the fixed frame. The vertical cylinders have no pivots; there is a perforation through the center of the lower or larger brass plate which, when the cylinder is mounted, is fitted upon a vertical spindle projecting upwards from the center of a second horizontal brass plate; this second brass plate sustains the weight of the vertical cylinder and turns horizontally, being supported by three antifriction wheels (each in a vertical plane) carried by the fixed frame.

Uniform rotatory motion is given to the cylinders by the action of clock-work, or rather chronometer-work, regulated by either duplex-escapement or chronometer-escapement. For two of the cylinders, which revolve in 24 hours, and for the thermometercylinder which revolves in 50 hours, the axis is placed in the center of the chronometer, and a fork at the end of the hour hand takes hold of a winch fixed to the plate of the cylinder, or (in the vertical cylinders) to the plate that sustains the cylinder. In the cylinder for galvanic earth-currents only, the connexion is made by toothed wheels. For the horizontal cylinders, the plane of the chronometer work is vertical; for the vertical cylinders, it is horizontal.

Three of the cylinders are $11\frac{1}{2}$ inches high, $14\frac{1}{4}$ inches in circumference; that for the thermometers is 10 inches high, and 19 inches in circumference.

Each cy1inder is coycred, when in use, by a tube of glass, which is open at one end,

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and has at the other end a circular plate of ebonite or brass, perforated at its center. The tube is a little larger than the cylinder; its open end is kept in position by a narrow collar of ebonite, and the opposite end by a circular piece of brass fixed to the smaller brass plate at the end of the cylinder.

To prepare the cylinder for register of indications, it is covered with a sheet of photographic paper; the moisture on the paper usually agglutinates its overlapping ends with sufficient firmness; the glass tube is then slipped over it, and the cylinder thus loaded is placed (if horizontal,) with its pivots in bearing upon its two sets of antifriction wheels, or, (if vertical,) with its end-brass-plate upon the rotating brass plate, and its central perforation upon the spindle of that plate; care is taken to ensure connection with the clock-work, and the apparatus is ready for action.

The light, by which the trace of each instrument is made, originates in a lamp, formerly of camphine, but, since 1849, of coal gas, sometimes charged with the vapour of coal-naphtha. Before the flame of the lamp is placed a metallic plate, with a small aperture about $0^{\text{in.3}}$ high and $0^{\text{in.1}}$ broad, independent of the lamp, and supported (for the magnetometers) by a part of the stone capping of the brick pier which carries the magnet; or (for the earth-current apparatus and thermometers) by the upper platform of the braced frame which carries the rest of the apparatus. The following arrangements are for the purpose of throwing on the photographic paper of the revolving cylinder a spot of light which shall travel in the direction of the cylinder's axis with every motion of either magnetometer, or of either galvanometer, or with the rise or fall of the mercury of the barometer or of either thermometer.

For each of the three magnetometers, a large concave mirror of speculum metal is carried by a part of the magnet-carrier; although it has a small movement of adjustment relative to the magnet-carrier, yet in practice it is very firmly clamped to it, so that the mirror receives all the angular movements of the magnet. The lalnp above mentioned is placed slightly out of the direction of the straight line drawn from the center of the concave mirror to the center of the cylinder which carries the photographic paper. By the concave mirror, the light diverging from the aperture is made to converge to a place nearly on the surface of the cylinder of photographic paper. The form of the aperture, however, and the astigmatism caused by the inclined reflexion from the mirror, produce this effect, that the image is somewhat elongated in the vertical direction, and is at the same time slightly curved. To diminish the length there is placed near the cylinder a plano-convex cylindrical lens of glass, with its axis parallel to the axis of the cylinder, and the image is thus reduced to a neat spot of light.

For the registers of galvanic earth-currents, the light, which falls upon a plane mirror carried by each galvanometer, is made to converge to a spot by a system of cylindrical lenses.

For the barometer, the light shines through a small aperture in a plate of blackened mica, which moves with the fluctuations of the quicksilver, and thus forms a spot of light.

\int GENERAL PRINCIPLE OF PHOTOGRAPHIC REGISTRATION. xv

For the thermometers, the light shines through the vacant part of the tube, and thus forms a sheet of light.

The spot of light (for the magnets, the earth-currents, and the barometer) or the boundary of the line of light (for the thermometers) moves, with the movements which are to be registered, in the direction of the axis of the cylinder, while the cylinder itself is turned round. Consequently, when the paper is unwrapped from its cylindrical form, there is traced upon it (though not visible till the proper chemical agents have been applied) a curve, of which the abscissa measured in the direction of a line surrounding the cylinder is proportional to the time, while the ordinate measured in the direction parallel 'to the axis of the cylinder is proportional to the movement which is the subject of measure.

In the instruments for registering the motions of the magnets, the earth-currents, and the barometer, a line of absciss to actually traced on the paper, by a lamp giving a spot of light in an invariable position, the effect of which on the revolving paper is to 'trace 'a line surrounding the cylinder. For the thermometers this is not necessary, as the thermometer-scales are made to carry and to transfer to the photographic paper sufficient indications of the actual reading of the thermometers.

Every part of the cylinder-apparatus for the declination and horizontal force, except those on which the spots of light fall, is covered with a double case of blackened zinc, having a slit for each moveable spot of light and a hole for the invariable spot; and every part of the path of the photographic light is protected by blackened zinc tubes from the admixture of extraneous light. The cylinder-apparatus for the thermometers is protected in the same manner, except that the whole space including the gas-light is enclosed in a zinc case, blackened internally. The earth-current apparatus is enclosed in a mahogany case, similarly blackened.

In all the instruments, the following method is used for attaching, to the sheet of photographic paper, indications of the time when certain parts of the photographic trace were actually made, and for giving the means of laying down a time· scale applicable to every part of the trace. By means of a smal1 moveable plate, arranged expressly for this purpose, the light which makes the trace can at any moment be completely cut off. An assistant, therefore, occasionally cuts off the light (registering in the proper book the clock-time of doing so), and after a few minutes withdraws the plate (again registering the time). The effect of this is to make a visible interruption in the trace, corresponding to registered times. By drawing lines from these points of interruption parallel to the axis of the cylinder, to meet the photographic line of abscissæ, or an adopted line of abscissæ parallel to it, points are defined upon the line of absciss a corresponding to registered times. The whole length of the photographic sheet (except where one end, in the cylindrical arrangement, laps over the other) corresponds to the known time of revolution of the cylinder. A scale being prepared 'beforehand, whose value for the time of revolution corresponds to the circumference of the cylinder, and the scale-reading for the registered time of interruption of light

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being applied to the foot of the ordinate corresponding to that interruption, the divisions of hours and minutes may be transferred at once from the scale to the line of abscissæ. In practice it is found that the length of the paper is not always the same, and it is necessary, therefore, to use for each instrument several pasteboard scales of different lengths, adapted to various lengths of the photographic sheets. ,

In the present year (1870), an opening has been made in the chimney of each of the lamps of the concave mirror; and the light in each instance falls upon the cylindrical lens, and produces a dark line upon the photographic paper. An apparatus of clockwork, specially arranged by Messrs. E. Dent and Co. for this purpose, uncovers simultaneously the chimney-holes in all the lamps about $2\frac{1}{2}$ minutes before each hour, and covers them all simultaneously about $2\frac{1}{2}$ minutes after each hour. In this manner a good series of hour-lines in the direction of the ordinates is formed. The system of cutting off the trace by hand is still retained, as giving means of correcting any error in the clock, $\&c.$; the correction thus found will be common to all the hour-lines. The accuracy of the time-registers has been much increased by this arrangement.

§ 4. *Lower Declination-Magnet; and Photographic self-registering Apparatus for Continuous Record* of *Magnetic Declination.*

The lower declination-magnet is made by Simms. It is 2 feet long, $1\frac{1}{2}$ inch broad, $\frac{1}{4}$ inch thick, of hard steel throughout, much harder than the upper declination-magnet.

The magnet-frame consists of an upper piece, whose top is a hook, (to be hooked into the suspension-skein), and which carries a concave mirror used for the photographic record in the manner described above. The lower part of this upper piece turns in a graduated horizontal circle, similar to the torsion circle of the upper magnet, and attached to the lower piece or magnet-carrier proper. The lowest part of the carrier is a double square hook, in which the magnet is inserted and is kept in position by the pressure of three screws.

It has been mentioned in \S 1 that a small pier, built upon one of the crossed slates which are laid upon three piers rising from below, carries the suspension-pulleys. The suspension-skein rises to one of these pulleys, passes horizontally over a second pulley about 5 inches south of it, and then descends obliquely to a windlass which is fixed to the stone slab about 2 ft. 3 in. south of the center of the magnet.

The height of the pulley above the floor of the Basement is 10 ft. 4²/₂ in. As the height of the magnet above the floor is 2 ft. $10\frac{1}{2}$ in., and the length of the magnet frame is 1 ft. 3 in., there remains 6 ft. $3\frac{1}{4}$ in. of free suspending skein.

On 1870, July 5, the skein, which had been in use from 1865, January 30, gave way. On July 7 a new skein was mounted; from July 7 to July 9 experiments were made for freeing it from torsion; and on July 9 the photographic registers were restored in their usual form.

One of the revolving cylinders is used for the photographic record of the Declination-Magnet and the Horizontal Force Magnet. In the preparation of the basement

· LOWER DECLINATION MAGNET: HORIZONTAL-FoRCE-MAGNET. *.xvii*

in 1864, as has been stated, the south-eastern re-entering angle was cut away, so that the straight line from the suspending skein of the declination-magnet to the center of those of the bifilar magnet passes through a clear space, in which the registering apparatus is placed.

The concave mirror of the declination-magnet is 5 inches in diameter, and is above the top of the magnet-box. The distance of the light-aperture from the mirror is about 25·3 inches. The bright spot formed by the reflection of light from the mirror is received on the south side of the cylinder, near its west end.

For the declination-magnet, the values, in minutes and seconds of arc, of movements of the photographic spot in the direction of the ordinate, are thus deduced from a geometrical calculation founded on the measures of different parts of the apparatus. The distance of the cylinder from the concave mirror is about 11^{ft} . 0^{in} . I, and a movement of 1° of the mirror produces a movement of 2° in the reflected ray. From this it is found that 1° of movement of the mirror is represented by 4.611 inches upon the photographic paper. A small scale of pasteboard is prepared, (for which a glass scale is now substituted), whose graduations correspond in value to minutes and seconds so calculated. The zero of the ordinate-scale is found in the following manner. The time-scale having been laid down as is already described, and actual observations of the position of the upper declination-magnet having been made with the eye and the telescope, (as has been fully described above), at certain registered times, there is no difficulty (by means of these registered times) in defining the points of the photographic trace which correspond to the observed positions. The pasteboard scale being applied as an ordinate to one of these points, and being slid up and down till the scale reading which represents the reading actually taken by the eye-observation fal1s on that point, the reading of the scale where it crosses the line of abscissæ is immediately found. This process rests on the assumption that the movements of the upper and lower magnets are exactly similar. The various readings given by different observations, so long as there is no instrumental change, will scarcely differ, and may be combined in groups, and thus an adopted reading for the line of abscissæ may be obtained. From this, with the assistance of the same pasteboard scale, there will be laid down without difficulty a new line, parallel to that line of abscissæ whose ordinate would represent some whole number of degrees, or other convenient quantity.

§ 5. *Horizontal-Foree-Magnet and Apparatus for observing it.*

The horizontal-force-magnet, furnished by Meyerstein of Göttingen, is, like the declination-magnet, 2 feet long, $1\frac{1}{2}$ inch broad, and about $\frac{1}{4}$ inch thick. For its support (as is mentioned above), a brick pier in the eastern arm of the Magnetic Observatory, built on the ground below the basement floor, rises through the floor of the upper room, and carries a slate slab, to the top of which a brass frame is attached,

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carrying two brass pulleys (with their axes in the same-east and west line) in front: of the pier, and two (in a similar position) at the back of the pier; these constitute: the upper suspension-piece. A small windlass is attached to the back of the pier at a convenient height. The magnet-carrier consists of two parts. The upper part is a horizontal bar, $2\frac{1}{2}$ inches long, whose ends are furnished with verniers for reading the graduations of the torsion-circle (a portion of the lower part, to be mentioned below). On the upper side of this horizontal bar are two small pulleys with axes horizontal and, at right angles to the vertical plane passing through the length of the bar: by these pulleys the apparatus is suspended, as will be mentioned. From the lower side of the horizontal bar, a vertical axis projects downwards through the center of the torsioncircle, in which it turns by stiff friction. The lower part of the magnet-carrier consists, first of the torsion-circle, a graduated circle about 3 inches in diameter: next, immediately below the central part of the torsion-circle, is attached (but not firmly fixed) a circular piece of metal from which projects downwards a frame that, by' means of three' cramps and screws, carries the photographic concave mirror, with the plane of its front under the center of the vertical axis: this circular piece of metal has a radial arm upon which acts a screw carried by the torsion-circle, for giving' to the concave mirror small changes of azimuthal position. Thirdly, there is fixed to the torsioncircle, at the back of the mirror-frame but not touching it,a bar projecting downwards, bent horizontally under the mirror-frame and then again bent downwards, carrying the cramps in which the magnet rests; and, still lower, a small plane mirror, to which a fixed telescope is directed for observing by reflexion the graduations of a fixed scale (to be mentioned shortly). Under the two small pulleys mentioned above passes a skein of silk; its two branches rise up and pass over the front pulleys of the suspension-piece, then over its back pulleys, and then descend and pass under a single large pulley, whose axis is attached to a wire that passes down to the windlass. Supported by the two branches of the skein, the magnet swings freely, but the direction that it takes will depend on the angular position of its stirrup with respect to the upper horizontal bar; it is intended that the index should be brought to such a position on the torsion-circle that the two suspending branches should not hang in one plane, but should be so twisted that their torsion-force will maintain the magnet in a direction very nearly E. and W. magnetic (its marked end being W.); in which state an increase of the earth's magnetic force draws the marked end towards the N., till the torsion-force is sufficiently increased to resist it; or a diminution allows the torsionforce to draw it towards the S. The magnet, with its plane mirror, hangs within a double rectangular box (one box completely inclosed within another) covered with gilt paper, similar to that used for the declination-magnet; in its S. side there is one long hole, covered with glass, through which the rays of light from the scale enter to fall on the plane mirror, and the rays reflected by the mirror pass to the fixed telescope. The vertical rod (below the torsion-circle), which carries the magnet-stirrup, passes through a hole in the top of the box. Above the magnet box is the concave mirror

above mentioned. The height of the brass pulleys of the suspension-piece above the floor is 11^{th} . 8^{in} . 5; that of the pulleys of the magnet-carrier is 4^{th} . 2^{in} . 5; and that of the center of the plane mirror is about 3^{th} lⁱⁿ. The distance between the branches of the silk skein, where they pass over the upper pulleys, is $1^{in.}14$; at the lower part the distance between them is $0ⁱⁿ·80$.

An oval copper bar (exactly similar to that for the declination-magnet), embraces the magnet, for the purpose of diminishing its vibrations.

The scale, which is observed by means of the plane mirror, is in a horizontal position, and is fixed to the South wall of the East arm of the Magnetic Basement. The numbers of the scale increase from, East to West, so that when the magnet *is* inserted in the magnet-cell with its marked end towards the West, increasing readings of the scale (as seen with a fixed telescope directed to the mirror which the magnet carries) denote an increasing horizontal force. A normal from the plane-mirror to the scale meets it at the division 51 nearly; the distance from the center of the plane-mirror to the scale is 7^{rt} . 6^{in} . \cdot 8.

The telescope is fixed on the east side of the brick pier which supports the stone pier of the declination-theodolite in the upper observing room. The angle between the normal to the scale (which usually coincides nearly with the normal to the axis of the magnet) and the axis of the telescope, is about 38°, and the plane of the mirror is therefore inclined to the axis of the magnet about 19° .

OBSERVATIONS RELATING TO THE PERMANENT ADJUSTMENTS OF THE HORIZONTAL-FORCE-MAGNET.

1. Determination of the times of vibration and of the different readings of the scale for different readings of the torsion-circle, and of the reading of the torsion-circle and the time of vibration when the magnet is transverse to the magnetic meridian .

. To render the process intelligible, it may be convenient to premise the following explanation.

Suppose that the magnet is suspended in its stirrup which is firmly connected with the small plane mirror, with its marked end in a magnetic westerly direction (not exactly W., but in any westerly direction between N. and S.), and suppose that, by means of the telescope directed towards that mirror, the scale is read, or (which is the same thing) the position of the plane mirror and of the stirrup, and therefore that of the axis of the magnet, are defined. Now let the magnet be taken out of the stirrup and replaced with its marked end easterly. The terrestrial magnetic power will now act as regards torsion, in the direction opposite to that in which it acted before, and

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therefore the magnet will not take the same position as before. But by turning the . torsion-circle, which changes the amount and direction of the torsion-power produced by the oblique tension of the suspending cords, the magnet may be made to take the same position as at first (which will be proved by the reading of the scale, as viewed in the plane mirror, being the same). The reading of the torsion-circle will be different from what it was. The effect of this operation then is, to give us the difference of torsion-circle-readings for the same position of the magnet-axis with the marked end opposite ways, but it gives no information as to whether the magnet-axis is accurately transverse to the meridian, inasmuch as the same operation can be performed whether the magnet-axis is transverse or not.

But there is another observation which will inform us whether the magnet-axis is or is not accurately transverse. Let the time of vibration be taken in each position of the magnet. Resolve the terrestrial magnetic force acting on the poles of the magnet into two parts, one transverse to the magnet, the other longitudinal. In the two positions of the magnet (marked end westerly and marked end easterly, with axis in the same position), the magnitude of the transversal force is the same, and the changes which the torsion undergoes in a vibration of given extent are the same, and the time of vibration (if there were no other force) would be the same. But there is another force, namely, the longitudinal force; and when the marked end is northerly, this tends from the center of the magnet's length, and when it is southerly it tends towards the center of the magnet's length; and in a vibration of given extent this produces force, in one case increasing that from the torsion and in the other case diminishing it. The times of vibration therefore will be different. There is only one exception to this, which is when the magnet-axis is transverse to the magnetic meridian, in which case the longitudinal force vanishes.

The criterion then of the position truly transverse to the meridian (which position is necessary in order that the indications of our instrument may apply truly to changes of the magnitude of terrestrial magnetic force without regard to changes of direction) is this. Find the readings of the torsion-circle which, with magnet in reversed positions, will give the same readings of the scale as viewed by reflexion in the plane mirror, and will also give the same time of vibration for the magnet. With these readings of the torsion-circle the magnet is transverse to the meridian; and the difference of the readings of the torsion-circle is the difference between the position when terrestrial magnetism acting on the magnet twists it one way, and the position when the same force twists it the opposite way, and is therefore double the angle due to the torsionforce of the suspending lines when they neutralize the force of terrestrial magnetism.

The following table exhibits the clements of one of the determinations made for $1870 :=$

The times of vibration and scale readings were sensibly the same, when the torsioncircle read 145°.30', marked end West, and 227°.30', marked end East, differing 82°. 0'. Half this difference, or 41° .0', is the angle of torsion when the magnet is transverse to the meridian.

The mean of several similar determinations gave 40° . $55^{\prime}\cdot0$; and this value was adopted in the reduction of observations through the year 1870.

The reading adopted for the torsion-circle, marked end of magnet west, was 145°.30' through the year. .

2. Computation of the angle corresponding to one division of the scale, and of the variation of the horizontal force (in terms of the whole horizontal force) which moves the magnet through a space corresponding to one division of the scale.

It was found by accurate measurements, on 1864, November 3, that the distance from 51^{div} on the scale to the center of the face of the plane mirror is 7^{ft} , $6ⁱⁿ$, 84, and that the length of 30div·85 of the scale is exactly 12 inches; consequently the angle at the mirror subtended by one division of the scale is 14'. 43"25, or, for one division of the scale, the magnet is turned through an arc of 7'. 21"·625.

The variation of horizontal force (in terms of the whole horizontal force) for a disturbance through one division of the scale, is computed by the formula, "Cotan. angle of torsion \times value of one division in terms of radius." Using the numbers of the last article, the value is found to be 0'00247025 through the year 1870.

3. Determination of the compound effect of the vertical-foree-magnet and the decli. nation-magnet on the horizontal-foree-magnet, when suspended with its marked end towards the West.

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The details of the experiments, made while the old vertical-foree-magnet was in use, will be found in the volumes for 1841, 1842, 1843, 1844, 1845. The effect was to increase the readings by $0^{div-487}$. On mounting a new vertical-force-magnet in 1848, similar experiments were made, and the resulting number was 0^{div-45} . These quantities are totally unimportant in their influence on the registers of changes of horizontal force. No experiments have been made since the magnets were placed in the basement.

4. Effect of the damper.

In the year 1865, from May 17 to May 25, observations were made for ascertaining the deflection of the magnet produced by turning the damper' through a small angle round n vertical axis passing through its center.

DAMPER IN USUAL POSITION.

On 1865, July 25, observations were made to ascertain whether the effect of an external deflecting cause is the same with the damper present and the damper removed. *A* small magnet was placed with its marked end pointing N. at the distance 4 feet S. of the unmarked end of the horizontal-force-magnet, deflecting the magnet through 1^{div.} of the scale, and the scale-readings were observed with the damper in its usual place and the damper away. Three experiments were made, containing twenty-four observations of position. Not the smallest difference of position of the horizontal-forcemagnet was produced by the presence or absence of the damper. The observations were very easy, and the result is certain.

No experiments on the damper have been made since 1865.

5. Determination of the correction for the effect of temperature on the horizontalforce-magnet.

In the Introduction to the volume of Magnetical and Meteorological Observations for 1847 will be found a detailed account of observations made in the years 1846 and 1847 for determination of this element. The principle adopted was that of observing the deflection which the magnet (to be tried) produces on another magnet; the magnet (to be tried) being carried by the same frame which carries the telescope that is directed to the plane mirror attached to the other magnet, and which also carries

ADJUSTMENTS, AND TEMPERATURE CORRECTION OF THE HORIZONTAL-FORCE-MAGNET. xxiii

the scale that is viewed in these experiments by reflection in that plane mirror. The rotation of'the frame was measured by a graduated circle about 23 inches in diameter. The magnet (to be tried) was always on the eastern side of the other magnet. It was enclosed in a copper trough, which was filled with water at different temperatures. One end of the magnet (to be tried) was directed towards the other magnet. The values found for correction of the results as to horizontal force determined with the magnet at temperature t° in order to reduce them to what they would have been if the temperature of the magnet had been 32° , expressed as multiples of the whole horizontal force, were, $*$

When the marked end of the magnet (to be tried) was West,

 $0.00007137 (t-32) + 0.000000898 (t-32)^{2}$.

,;

When the marked end of the magnet (to be tried) was East,

 0.00009050 $(t-32)$ + 0.000000626 $(t-32)^2$.

The mean, or

 $0.00008093 (t-32) + 0.000000762 (t-32)^2$

has been embodied in tables which have been used in the computation of the " Reduction of Magnetic Observations 1848-1857," attached to the Volume of Observations 1859, and in the computation for "Days of Great Magnetic Disturbance 1841-1857," attached, to the volume for 1862. The same formula has been employed in the Reduction of Magnetic Observations 1858-1863, published in the volume for 1867.

. In the year 1864 observations were made for ascertaining the temperature-coefficient by heating the magnet by hot air. The magnet, whose variation of power in different temperatures was:to be determined, was placed in a copper box planted upon the top of a copper gas~stove, whose heat could be regulated by manipulation of a tap, and from which rose a stream of heated air (not the air vitiated by combustion) through a large opening in the bottom of the box. The stove used for this purpose was the same which is now used for warming the Magnetic Basement. It was placed in the Magnetic Office, No. 7, in a position magnetic south of the deflexion-apparatus used in the operation for ascertaining the absolute measure of horizontal magnetic force. The hot air which rose through the opening in the center of the bottom was discharged by; adjustible openings near the extreme ends of the top. Three windows were provided for reading three thermometers. The box, and the magnet which it inclosed, were placed in a, magnetic E. and W. position.' The needle whose deflection exhibited the power of the magnet was that which is employed in the ordinary use of the deflexionapparatus. The proportion of the power of the magnet (under definite circumstances) to the earth's directive horizontal power was expressed by the tangent of the angle of deviation. Observations were made with temperatures both ascending and descending.

^{*} By inadvertence in printing the Introduction 1847, the letter t has been used in two different senses.
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The intervals of observation at different temperatures were sufficiently small to permit the assumption that the earth's force had not sensibly changed. The following is an abstract of the principal results:-

Omitting some days of less perfect series, satisfactory series of observations were made on 1864, February 21, 22, 23, and March 10. The tangents of angle of deflection were as follows :-

From these it was inferred that the tangent of angle. of deflection could be represented by-

0.404559
$$
\times
$$
 { 1 - 0.0004610 \times (t - 32) + 0.000005061 \times (t - 32)² }

On comparing the quantity within the bracket (which expresses the law of magnetic power as depending on temperature) with that found in 1847, which, as above stated, is $-$

$$
\left\{1 - 0.00008093 \times (t - 32) - 0.000000762 \times (t - 32)^2\right\}
$$

it will be seen that the difference is great. The second terms differ greatly in magnitude, and the third terms in sign.

Possibly some light may be thrown on the difference by the following remark. The two formulæ give the same values for $t = 32^{\circ}$ and for $t = 97^{\circ}$. And they give equal degrees of change per degree when $t = 65^{\circ}$. It would seem therefore that the real discordance is in the experimental values for the mean temperatures only, or principally; and that it is probable that there is some error in the hot-air process for the middle temperatures.

I insert here (although not applying to the observations of the present volume) the results of a similar examination of the Old Vertical Force Magnet, which was in use to the end of 1863. Omitting less perfect series, observations made on 1864, February 21 and 24, gave the following values for tangents of angles of deflection:-

From these it was inferred that the tangent of angle of deflection could be represented by-

$$
0.280526 \times \left\{ 1 - 0.00088607 \times (t - 32) + 0.0000045594 \times (t - 32)^2 \right\}
$$

The expression found in 1847 for the law of force was-

 $\left\{1 - 0.00015816 \times (t - 32) - 0.000001172 \times (t - 32)^2\right\}$

giving a discordance of the same kind as that found for the horizontal force, but still larger. The formulæ agree only when $t = 32^{\circ}$ and when $t = 159^{\circ}$. The discordance cannot be removed by a supposition similar to that made above.

Returning now to the temperature-correction of the Horizontal Force Magnet. The unsatisfactory character of the comparisons just given induced me at the beginning of 1868 to try the method of heating the air of the Magnetic Basement generally (by means of the gas-stove), leaving the magnets in all respects in their ordinary state, and comparing their indications as recorded in the ordinary way, but at different temperatures.* Experiments were at first made at intervals of a few hours in the course of one day, but it was soon found that the magnet did not acquire the proper temperature; moreover, the result was evidently affected by diurnal inequality. After this, an entire day was in each case devoted to the effects of each temperature (high or low, as the case might be). The principal series of observations were made with the horizontal force magnet in its ordinary position, or marked end to the west; but a few were made with the marked end to the east. In some instances, the numbers given are the result each of several observations; but in other instances, the result is that of a single observation, taken when all the apparatus had acquired unusual steadiness. The following are the results :-

1868. MONTH and DAY. (Civil.)		Temperature.	Scale Reading.	Change оf Temperature.	Change of Scale Reading.	Change of Scale Reading reduced to Parts of the whole Horizontal Force.	Change of H.F. corresponding to a change of 1° of Temperature (in Parts of the whole Horizontal Force).	
		\bullet	div.	\circ	div.			
January	3 \mathbf{A}	56.8 50.5	60.82 61.47	6.3	\circ 65	0'001579	0.000250	
	4 4	49.5 55.5	61.47 61.35	6.0	0'12	'000202	000049	
	6 7 9	59.3 49.3 $5\overline{6}\cdot 7$	60.91 61.62 61.05	10.0 7.4	0.71 0.57	\cdot 001725 °001385	'000172 000187	
	10 \mathbf{I} 12	58.9 51.3 59.3	60.91 61.71 61:18	7.6 8.0	0.80 0.53	\cdot 001943 \cdot 001288	•ooo256 .0001Q1	

RESULTS OF TEMPERATURE EXPERIMENTS UPON THE H,F. MAGNET MARKED END WEST.

* This method was first used for magnets, so far as I am aware, at the Kew Observatory. It had been used for pendulums by Lieut,-General Sir Edward Sabine and by myself.

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 \mathbf{P}^{in} , and the state of the state of the state of \mathbf{P}

 $\mathcal{L}^{\text{max}}_{\text{max}}$ and $\mathcal{L}^{\text{max}}_{\text{max}}$

 $\sim 10^{-11}$

RESULTS OF TEMPERATURE EXPERIMENTS UPON THE H.F. MAGNET

MARKED END *WEST-continued.*

RESULTS OF TEMPERATURE EXPERIMENTS UPON THE H.F. MAGNET MARKED END EAST.

 \sim

EYE-OBSERVATIONS, AND PHOTOGRAPHIC APPARATUS OF THE HORIZONTAL-}'OROE-MAGNET. *xxvii*

These results do not differ greatly from those which are given by application of the formula found in 1847. It is important to observe that they include the entire effects of temperature upon all the various parts of the mounting of the magnet, as well as on the magnet itself; and for this reason I think them deserving of great confidence. Still I have thought it prudent, at present, to omit application of corrections for temperature.

The method of observing with the horizontal-force-magnet is the following: $-\frac{1}{2}$

A fine vertical wire is fixed in the field of view of the telescope, which is directed to the plane mirror carried by the magnet. On looking into the telescope, the graduations of the fixed scale, mentioned in pages *xviii* and *xix*, are seen; and during the oscillations of the magnet, the divisions of the scale are seen to pass alternately right and left across the wire. The clock-time, for which the position of the magnet is to be determined, is the same as that for the observation of declination. The first observation is made by the observer applying his eye to the telescope 40^s before that time, and, if the magnet is in a state of vibration, heohserves the next four extreme points of vibration of the scale, and the mean of these is adopted in the same manner as for the declinationobservations; but if it is at rest, then at 10^s before the pre-arranged time, he notes the division of the scale bisected by the wire; and 10^s after the pre-arranged time he notes whether the same division continues bisected, and if it does, that reading is adopted as the result.

The number of instances when the magnet was observed in a state of vibration during the year 1870 is very small.·

Outside the double box is suspended a thermometer which is read on every day except Sundays, at 21^h , 22^h , 23^h , 0^h , 1^h , 2^h , 3^h , and 9^h . Occasional observations have been taken at other hours. Self-registering maximum and minimum thermometers placed outside the box were read twice every-day, but in consequence of the very small diurnal range of temperature, their readings are not printed in the volume.

§ 6. Photographic self-registering Apparatus for Continuous Record of Magnetic *Horizontal Force.*

Referring to the general description of photographic apparatus, the following remarks apply more particularly to that which is attached to the horizontal-foree-magnet. A concave mirror of speculum-metal, 4 inches in diameter, is carried by the magnet-carrier. The light of a gas-lamp shines through a small aperture 0^{in} 3 high, and 0^{in} 01 broad (which is supported by the solid base of the brick pier carrying the magnetsupport), at the distance of about 21.25 inches from the concave mirror, and is made to

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converge to a point, on the north surface and near the east end of the same revolving cylinder which receives the light from the concave mirror of the declination~magnet. A cylindrical lens parallel to the axis of the cylinder receives the somewhat elongated image of the source of light, and converts it into a well-defined spot. The motions of this spot parallel to the axis represent the angular movements of the magnet which are produced by an increase of terrestrial magnetic force overcoming more completely the torsion-force of the bifilar suspension, or by a diminution of terrestrial force yielding to the torsion-force.

As the spot of light from the horizontal-force-mirror falls on the side of the cylinder opposite to that on which the light from the declination-mirror falls, the same timescale will not apply to both; it is necessary to prepare a time-scale independently for each.

The following is the calculation by which the scale of horizontal force on the photographic sheet is determined. The distance between the surface of the concave mirror and the surface of the cylinder is 134'436 inches; consequently, one degree of angular motion of the magnet, producing two degrees of angular motion of the reflected ray, moves the spot of light through 4'6927 inches. For the year 1870 the adopted value of variation of horizontal force for one degree of angular motion of the magnet is sin $1^{\circ} \times \text{cotan } 40^{\circ}$. 55' = 0.020137; and the movement of the spot of light for 0'01 part of the whole horizontal force is 2'330 inches. With this fundamental number, the graduations of the pasteboard scale for measure of horizontal force have been prepared.

§ 7. *Vertical-Force-Magnet, and Apparatus for observing it.*

The vertical-force-magnet in use to 1848 was made by Robinson; that in use from 1848 to 1864, January 20, was by Barrow. The Inagnet now in use is by Simms. Its length is $1^{ft.}$ $6^{in.}$; it is pointed at the ends. After some trials, it was re-magnetized by Mr. Simms on 1864, June 15. Between 1864, August 27, and September 27, a new knife-edge was attached to it, to remedy a defect which, as was afterwards found, arose from a cause that had no relation to the knife-edge. Its supporting frame rests upon a solid pier, built of brick and capped with a thick block of Portland stone, in the western arm of the magnetic basement. Its position is as nearly as possible symmetrical with that of the horizontal-foree-magnet in the eastern arm. Upon the stone block is fixed the supporting frame, consisting of two pillars (connected at their bases) on whose tops are the agate planes upon which vibrate the extreme parts of the knife-edge (to be mentioned immediately). The carrier of the magnet is an iron frame, to which is attached, by clamps and pinching screws, a steel knife-edge, about 8 inches long. 'fhe steel knife-edge passes through an aperture in the magnet. The axis of the magnet is as nearly as possible transverse to the meridian,

HORIZONTAL FORCE PHOTOGRAPHY, AND VERTICAL FORCE-MAGNET. xxix

its marked end being E. The axis of vibration is as nearly as possible N. and S. To the southern end of the iron frame, and projecting further south than the end of the knife-edge, is fixed a small plane mirror, whose plane makes with the axis of the magnet an angle of $52\frac{3}{4}$ nearly. The fixed telescope (to be mentioned) is directed to this mirror, and by reflexion at the surface of the mirror it views a vertical scale (to be mentioned shortly). The height of this mirror above the floor is about $2^{rt} \cdot 10^{tn} \cdot 6$. Before the introduction of the photographic methods, the magnet was placed in a perforation of a brass frame midway between its knife-edges. But since the photographic method was introduced, the magnet has been placed excentrically; the distance of its southern face from the nearest end of the southern knife-edge being nearly 2 inches, and a space of $4\frac{1}{2}$ inches in the northern part of the iron frame being left disposable. In this disposable space there is attached to the iron frame by three clips a concave mirror of speculum-metal, with its face at right angles to the length of the magnet; it is used in the photographic system (shortly to be described). Near the north end of the iron frame are fixed in it two screw-stalks, upon which are adjustible screw-weights; one stalk is horizontal, and the movement of its weight affects the position of equilibrium of the magnet (which depends on the equilibrium between the moments of the vertical force of terrestrial magnetism on the one hand and of the magnet's center of gravity on the other hand); the other stalk is vertical, and the movement of its weight affects the delicacy of the balance, and varies the magnitude of its change of position produced by a change in the vertical force of terrestrial magnetism.

The whole is inclosed in a rectangular box. This box is based upon the stone block above mentioned; and in it, in a space separated from the rest by a thin partition, the magnet can vibrate freely in the vertical plane. In the south side of the box is a hole covered by glass, through which pass the rays of light from the scale to the plane mirror, and through which they are reflected from the plane mirror to the telescope. And at the east end is a large hole covered by glass, through which passes the light from the lamp to the concave mirror, and through which it is reflected to the photographic cylinder (to be described hereafter).

The telescope is fixed to the west side of the brick pier which supports the stone pier in the upper room carrying the declination-theodolite. Its position is symmetrical with that of the telescope by which the horizontal-force-magnet is observed; so that a person seated in a convenient position can, by an easy motion of the head left and right, observe the vertical-force and horizontal-foree-magnets.

The scale is vertical: it is fixed to the pier which carries the telescope, and is at a very small distance from the object-glass of the telescope. The wire in the field of view of the telescope is horizontal. The telescope being directed towards the mirror, the observer sees in it the divisions of the scale passing upwards and downwards over the fixed wire as the magnet vibrates. The numbers of the scale increase from top to

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bottom; so that, when the magnet is placed with its marked end towards the East, increasing readings (as seen with the fixed telescope) denote an increasing vertical force,

OBSERVATIONS RELATING TO THE PERMANENT ADJUSTMENTS OF THE VERTICAL-FORCE-MAGNET.

1. Determination of the compound effect of the declination-magnet, the horizontalforce-magnet, and the iron affixed to the electrometer pole, on the vertical-force. magnet.

The experiments applying to the magnets are given in the volumes for 1840-1841 to 1845: and those applying to the electrometer pole in the volume for 1842. It appeared that no sensible disturbance was produced on the Inagnet formerly in use. No experiments have been made with the new magnet.

2. Determination of the time of vibration of the vertical-foree-magnet in the vertical plane.

In the year 1870, vibrations of the vertical-force-magnet were observed on 164 different days, and with readings of various divisions of the scale. The mean time of vibration adopted for the year was 16^{s} -21.

3. Determination of the time of vibration of the vertical-force-magnet in the horizontal plane.

1868, December 31. The magnet with all its apparatus was suspended from a tripod in Magnetic Office, No.5, its broad side being in a plane parallel to the horizon; therefore, its moment of inertia was the same as when it is in observation. A telescope, with a wire in its focus, was directed to the reflector carried by the magnet. A scale of numbers was placed on the floor of the room, at right angles to the long axis of the magnet, or parallel to the mirror. The magnet was observed only at times when it was swinging through a small arc. From 500 vibrations, the mean time of one vibration $= 16°3192$. This number is used through the year 1870.

4. Computation of the angle through which the magnet moves for a change of one division of the scale; and calculation of the disturbing force producing a movement through one division, in terms of the whole vertical force.

The distance from the scale to the mirror is 186'07 inches, and each division of the scale $=\frac{12}{30.85}$ inches. Hence the angle which one division subtends, as seen from the mirror, is 7'. 11" 19; and therefore the angular movement of the normal to the mirror, corresponding to a change of one division of the scale, is half this quantity, or 3'. 35"·60.

But the angular movement of the normal to the mirror is not the same' as the angular movement of the magnet; but is less in the proportion of unity to the cosine

小野 むしょ ADJUSTMENTS OF VERTICAL-FORCE-MAGNET.

of the angle which the normal to the mirror makes with the magnet, or in the proportion of unity to the sine of the angle which the plane of the mirror makes with the magnet. This angle has been found to be $52\frac{3}{2}$; therefore, dividing the result just obtained by sine $52\frac{2}{4}$, we have, for the angular motion of the magnet corresponding to a change of one division of the scale, 4'. 30"'85.

From this, the value, in terms of the whole vertical force, of the disturbing force, producing. a change of one division, is to be computed by the formula, "Value of Division in terms of radius \times cotan dip $\times \frac{T^2}{T^2}$;" where T' is the time of vibration in the horizontal plane, and T the time of vibration in the vertical plane.

For the year 1870, T' was assumed = $16^{\circ}319$, T = $16^{\circ}21$, dip = 67° . 52'. 25". From these numbers, the change of the vertical force, in terms of the whole vertical force, corresponding to one division of the scale, is found $= 0.0005411$.

5. Investigation of the temperature-correction of the vertical-force-magnet.

The new vertical-force-magnet was subjected to experiments by inclosing it in a copper box, and warming it by an injection of hot air, and observing the amount of deviation which it produced on the suspended magnet used in the deflexion-apparatus for absolute measure of horizontal force, at the same time and in the same manner as were the horizontal-foree-magnet and the old vertical-foree-magnet, in the experirnents described in pages *xxiii* to *xxv*. Observations made on 1864, February 20, 25, March 3, 9, gave, for the tangents of the angles of deflection,—

From these it appeared that the angle of deflection might be represented by-

 $.0.172522 \times \left\{ 1 - 0.0002233 \times (t - 32) + 0.000001894 \times (t - 32)^2 \right\}$

The quantity within the brackets (which represents the variation of magnetic power in terms of the whole power of the magnet) shows the same peculiarities as those found for the other magnets; that the third term is large, and has a sign opposite to that of the second term.

The factor of variation for 1° of Fahrenheit, when $t = 62^{\circ}$, is $-$ 0'0001097.

. After these observations, the new vertical-foree-magnet was re-magnetized by Mr. Simms, on 1864, June 15.

In the beginning of 1868, observations were made in the method already described for the horizontal-foree-magnet, by heating the magnetic basement to different tempexxii INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

ratnres, and observing the scale-reading in the ordinary way. The results are as $follows :=$

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1868. MONTH and DAY.		Temperature.	Scale Reading.	Change of Temperature.	Change of Scale Reading.	Change of Scale Reading reduced to Parts of the whole Vertical Force.	Change of V.F. corresponding to a change of 1° of Temperature (in Parts of the whole V.F.)	
January	3 $\frac{4}{5}$	$56°$ o 48.2 5q.6	div. 56.45 46.52 61'49	\mathbf{o} 7.8 11'4	div. 9.93 14.97	0.006482 roog772	'000831 \cdot 000 857	
January February	6 7 10 11 12 13 14 16 17 18 20 22 2 ₃ 25 26 29 3 ₁ $\frac{4}{5}$ 6 7 8 10	59.6 49° 59.5 49.7 62.0 53.4 55.4 52.3 63.7 52.4 60.7 50.6 59.6 49.6 60.5 49.3 63∙ 1 51.0 62.3 50.6 53.3 50.6 $62 \cdot I$	61.73 46.84 61.62 48.70 64.40 53.33 55.72 50.79 $66 \cdot 13$ 53.26 62.19 47.82 5q 00 46.67 60.62 44.78 64.55 47'11 64.02 46.43 49.10 45.55 62.76	1016 10.5 9.8 12.3 8.6 2^{\degree} $3 \cdot I$ 11.4 11.3 8.3 10.1 $\mathbf{a} \cdot \mathbf{o}$ 10.0 10.0 11'2 13.8 12.1 $11 \cdot 3$ II.7 2.7 2.7 11.5	14.89 14.78 12.92 15.70 11'07 2.3a 4.93 15.34 12.87 8.93 14.37 11.78 12.03 13.95 15.84 19'77 17'44 16.91 17.59 2.67 $3 \cdot 55$ 17'21	0'009720 ∙oog648 .008434 010249 •007226 \cdot 001560 003218 • 010014 .008402 .005829 roog381 ∙oo⁊бgo .008441 '009107 0.0340 012906 011385 *01103q 0.11483 001743 002317 \cdot 011235	'000017 oooq1q •ооо861 \cdot 000 833 000840 .000780 001038 ro 00878 \cdot 000743 '000702 .000d5d \cdot coo 85 4 .000844 \cdot 000836 ooog23* \cdot 0000 35 ro ₀ q ₄ 1 '000977 000081 ∙000б4б ∙ooo858 ro ₀₉₇₇	
February	14 16 18	60.6 40.0 61.d	57.70 36.75 58.85	11.6 12.9	20.02 22'10	$'$ 011298 '011919	.000974 roo ₉₂₄	
February	18 20 21	61.9 50° 62.6	58:05 41.96 56.82	11.3 12.6	16.00 14.86	011749 010851'	.ooog87 \cdot 000 861	
Mean	\bullet		. .	\bullet	. .	\cdots	0.000880	

RESULTS OF TEMPERATURE EXPERIMENTS UPON THE VERTICAL-FORCE-MAGNET.

The coefficient of temperature-correction given by these experiments is enormously greater than any that has been found in any previous experiments. Yet I conceive that there can be no doubt of its accuracy. And it is easy to see that an instrument, subjected to the effects of gravity working differentially on its two ends, is liable to great changes depending on temperature which have no connexion with magnetism. For instance, if the point, at which the magnet is grasped by its carrier, is not absolutely coincident with its center of gravity, a great change of position may be produced by a small change of temperature. There appears to be no way of avoiding

TEMPERATURE COEFFICIENT: EYE-OBSERVATIONS: AND **PHOTOGRAPHIC APPARATUS OF THE VERTICAL-FORCE-MAGNET.** *xxxiii* x

these evils but by maintaining almost uniform temperature; a condition which has been almost perfectly preserved in the year 1870.

The method of observing with the vertical-force-magnet is the following $:$

A fine horizontal wire is fixed in the field of view of the telescope, which is directed to the small plane mirror carried by the magnet. On looking into the telescope, the graduations of the fixed vertical scale are seen; and during the oscillations of the magnet, the divisions of the scale are seen to pass alternately upwards and, downwards across the wire. The clock-time, for which the position of the magnet is to be determined, is the same as that for the other two magnets. The observer applies his eye to the telescope about two vibrations before the arranged time, and if the magnet is in motion he observes its places at four extreme vibrations; and the mean of these is taken as for the horizontal-foree-magnet. But if the magnet is at rest, then at one-half time of vibration before the arranged time, and at an equal interval after the arranged time, the division of the scale is noted; if there is a slight difference, the mean is taken.

The number of instances in 1870 in which the magnet was found in a state of vibration is very small.

Outside the box is placed a thermometer, which is read on every day except Sundays, at the hours 21^{h} , 22^{h} , 23^{h} , 0^{h} , 1^{h} , 2^{h} , 3^{h} , and 9^{h} . Occasional readings of the thermometer are also taken at other hours.

A maximum and a minimum thermometer have also been read twice daily; but the results are not printed.

§ 8. Photographic self-registering Apparatus for Continuous Record of Magnetic *Vertical Force.*

The concave mirror which is carried by the vertical-force-magnet is 4 inches in diameter; its mounting has been described in the last article. At the distance of about 22 inches from that mirror, and external to the box, is the horizontal aperture, about $0ⁱⁿ$ 3 in length and $0ⁱⁿ$ 01 in breadth, carried by the same stone block which carries the supports of the agate planes. The lamp which shines through this aperture is carried by a wooden stand. The light reflected from the mirror passes through a cylindrical lens with its axis vertical, very near to the cylinder currying the photographic paper, and finally forms a well-defined spot of light on the cylinder of paper, at the distance of 100'18 inches from the mirror. As the movements of the magnet are vertical, the axis of the cylinder is vertical. The cylinder is about $14\frac{1}{4}$ inches in circumference, being of the same dimensions as those used for the declination and horizontal-force magnets, and for the earth-currents. The forms of the exterior and interior cylinders, and the method of mounting the paper, are in all respects the same as for the declination and horizontalforce magnets; but the cylinder is supported by being merely planted upon a circular

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horizontal plate (its position being defined by fitting a central 'hole in the metallic cap of the cylinder upon a central pin in the plate), which rests on anti-friction rollers and is turned by watchwork once in twenty-four hours. The trace of the vertical-forcemagnet is on the west side of the cylinder.

On the east side, the cylinder receives the trace produced: by the barometer (to be described hereafter). A pencil of light from the lamp which is used for the barometer shines through a fixed aperture with a small cylindrical lens, for tracing a photographic base-line upon the cylinder of paper, similar to that for the cylinder of the declination and horizontal-force magnets.

The scale for the ordinates of the photographic curve of the vertical force is thus computed. Remarking that the radius which determines the range of the motion of the spot of light is double the distance 100 \cdot 18 inches, and is therefore $= 200 \cdot 36$ inches, the formula used in the last section, when applied to $\frac{\text{distribing force}}{\text{whole vertical force}} = 0.01$, gives value of division = 200.36 \times tan. dip. $\times (\frac{T}{T})^2 \times 0.01$. The value of the ordinate of the photographic curve for $\frac{\text{distribing force}}{\text{whole vertical force}} = 0.01$, thus obtained, is, for the year $1870 = 4.862$ inches. With this value, the pasteboard scales, used for measuring the photographic ordinates, have been prepared.

§ 9. *Dippin{? Needles, and Method of observing the Magnetic Dip. '*

The instrument with which the greater number of the dips in the year 1870 have been observed, is that which, for distinction, is called Airy's instrument. The following description will probably suffice to convey an idea of its peculiarities :—

The form of the needles, the form of their axes, the form of the agate bearings, and the general arrangement of the relieving apparatus, are precisely the same as those in Robinson's and other needles. But the form of the observing apparatus is greatly modified, in order to secure the following objects: $-$

I. To obtain a microscopic view of the points of the needles, as in the instruments introduced by Dr. Lloyd and Lieut.-General Sir E. Sabine.

II. To possess at the same time the means of observing the needles while in a state of vibration.

III. To have the means of observing needles of different lengths.

IV. To give an illumination to the field of view of each microscope, directed from the side opposite to the observer's eye, so that the light may enter past the point of the needle into the object glass of the microscope, forming a black image of the needlepoint in a bright field of view.

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'T. To give facility for observing by day or night.

$\label{eq:2} \mathcal{L}^{\mathcal{A}}(\mathcal{A},\mathcal{B})=\mathcal{L}^{\mathcal{A}}(\mathcal{A},\mathcal{B})=\mathcal{L}^{\mathcal{A}}(\mathcal{A},\mathcal{B},\mathcal{B})=\mathcal{L}^{\mathcal{A}}(\mathcal{A},\mathcal{B})$ DIP INSTRUMENT.

With these views, the following form is given to the apparatus $:$

The needle, and the bodies of the microscopes, are inclosed in a square box. The base of the box, two vertical sides, and the top, are made of gun-metal (carefully selected to insure its freedom from iron ; but the sides parallel to the plane of vibration of the needle are of glass. Of the two glass sides, that which is next the observer is firmly fixed; it is hereafter called " the graduated glass-plate." The other glass side can be withdrawn, to open the box, for inserting the needle, &c.

An axis, whose length is perpendicular to the plane of vibration of the needles, and is as nearly as possible in the line of the axis of the needle, supported on two bearings (of which one is cemented in a hole. in the graduated glass-plate, the other being upon a horizontal bar near to the agate support of the needle-axis), carries a transverse arm, about 11 inches long, or rather two arms, projecting about $5\frac{1}{2}$ inches on each side of the axis. Each of these projecting arms originally had a long opening, or slot, about 1 inch wide, extending from the neighbourhood of the center-work nearly to the end of the arm. Through this opening the tube of a microscope passed, in a direction parallel to the axis of the needle, and was firmly fixed by a shoulder-bearing on one side of the arm, and a circular nut, working in a thread cut upon the microscope-tube, on the other side of the arm. The microscope could thus be fixed at any distance from the central axis, within the limits of the length of the projecting arm. In 1863, between February 24 and May 11, the slot for a single moveable microscope on each side was changed for three fixed microscopes on each side, adapted in position to the lengths of the needles to be mentioned shortly.

The microscope-tube thus carried is not the entire microscope, but so much as contains the object-glass and the field-glass. Upon the plane side of the field-glass (which is turned towards the object-glass), a series of parallel lines is engraved by etching with fluoric acid. The object-glass is so adjusted that the image of the needle-point is formed upon the plane side of the field-glass; and thus the parallel lines can be used for observing the needle in a state of vibration; and, one of them being adopted as standard, the lines can be used for reference to the graduated circle (to be mentioned). All this requires that there be an eye-glass also for the microscope.

The axis of which we have spoken is continued through the graduated glass-plate, and there it carries another transverse arm parallel to the former, and generally similar to it. In each part of this there was originally a sliding eye-socket carrying the eyeglass. In 1863, at the time mentioned above, the slotted arm and moveable eyesocket were changed for an arm with three sockets and eye-glasses. Thus, reckoning from the observer's eye, there are the following parts:-

(1.) The eye-glass.

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(2.) The graduated glass-plate (its graduations, however, not intervening in this part of the glass, the graduated circle being so large as to include all the microscopes). (3.) The field-glass, on the further surface of which the parallel lines are engraved.

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(4.) The object-glass.

(5.) The needle.

(6.) The removeable glass side of the box.

(7.) The illuminating reflector, to be described hereafter.

The optical part of the apparatus being thus described, we may proceed to speak of the graduated circle.

The graduations of the circle (whose diameter is about $9\frac{3}{2}$ inches) are etched on the inner surface of the graduated. glass-plate. These divisions (as well as the parallel lines on the field glasses of the microscopes) are beautifully neat and regular, and are, I think, superior to any that I have seen on metal. The same piece of metal, which carries the transverse arms supporting the microscope bodies, carries also two arms with verniers for reading their graduations. These verniers (being adapted to transmitted light) are thin plates of metal, with notches instead of lines. The reading of the verniers is very easy. The portion of the axis which is external to the graduated glass-plate (towards the observer), and which has there, as already stated, two arms for carrying the microscope eye-glasses, has also two arms for carrying the lenses by which the verniers and glass-plate graduations are viewed. These four arms are the radii of a circle, which can be fixed in position by a clamp, attached to the gun-metal casing of the graduated glass-plate, and furnished with the usual slow-motion screw.

The entire system of the two arms carrying the microscope-bodies, the two arms carrying the microscope eye-glasses, the two arms carrying the verniers, and the two arms carrying the reading-glasses for the verniers, is turned rapidly by means of a button on the external side of the graduated glass-plate, or is moved slowly by means of the slow-motion screw just mentioned.

It now remains only to describe the illuminating apparatus. On the outside of the removeable glass plate, there are supports for the axis of a metallic circle turning in a plane parallel to the plane of needle-vibration. This circle has four slotted radii, and in these slots or openings there slide small frames carrying prismatic glass reflectors, each of which can turn on an axis, in the plane of the circle but transverse to the radius. Two of these reflectors are for the purpose of sending light through the verniers, and therefore ure fixed in radial distance; the other two were intended for sending light past the ends of the needle through the microscopes, and therefore required adjustment on change of needle and corresponding change of position of microscopes. In 1863 these were changed for fixed reflectors, corresponding to the fixed microscopes. 'fhe circle was originally turned by a small winch near the observer's hand; at present, the winch is removed, as its axis was found to be slightly magnetic. At each observation, it is necessary to turn the circle which carries the reflectors; but this is the work of an instant.

The light which illuminates the whole is a gas-burner, in the line of the axis of rotation. Its rays fall upon the glass prisms, and each of these is adjusted, by turning on its axis, to throw the reflected light in the required direction.

DIP INSTRUMENT: ABSOLUTE MEASURE OF HORIZONTAL MAGNETIO FORCE. *xxxvii*

The whole of the apparatus, as thus described, is planted upon a horizontal plate admitting of rotation in azimuth: the plate is graduated in azimuth, and verniers are fixed to the gun-metal tripod stand. The gas-pipe is led down the central vertical axis, and there communicates by a rotatory joint with the fixed gas-pipes.

The needles adapted for use with this instrument are—

The needles constantly employed are B_1 , C_1 , D_1 , B_2 , C_2 , D_2 .

In discussing carefully the observations taken with this instrument (as well as with other dip-instruments), great trouble was 'experienced in determining the zenith-point (or reading of the vertical circle when the points of the needle are in the same vertical). 'To remedy this, a "zenith-point-needle" was constructed under my instructions by Mr. Simms; and it has since been used as need required. It is a flat bar of brass; with pivots similar to those of the dip-needles; and with three pairs of points corresponding to the three lengths of needles used; loaded at one end so as to take a position perfectly definite with respect to the direction of gravity; observed with the microscopes, and reversed for another observation, exactly as the dip-needles. For each of the different lengths of dip-needles, the zenith-point is determined by observation of that pair of points of the zenith-point-needle whose interval is the same as the length of the dip.needle.

The Dip Instrument and all the needles are examined, at the close of each year and at other times if thought desirable, by Mr. Simms.

Besides this instrument, which is the property of the Royal Observatory, two dipinstruments of the Kew pattern were also used in 1870. One by Dover was used several times in the month of January; the other by Simms was occasionally used from January to October.

§ 10. *Observations for the absolute Measure* of *tIle Horizontal Force* of *Terrestrial Magnetism.*

In the spring of 1861, a Unifilar Instrument, similar in all respects (as is understood) to those used in and issued by the Kew Observatory, was procured by the xxxviii INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

courteous application of Sir E. Sabine, from the makers, Messrs. J. T. Gibson and Son; and after having been subjected to the usual examinations, at the Kew Observatory, for determination of its constants (for which I am indebted to the kindness of Balfour Stewart, Esq.), was mounted at the Royal Observatory. Observations with this instrument commenced on 1861, June 11, and were continued through the year; and, after some slight modifications of its verniers, it is still maintained in use (1870).'

The deflected magnet (whose use is merely to ascertain the proportion which the power of the deflecting magnet at a given distance bears to the power of terrestrial magnetism) is 3 inches long, carrying a small plane mirror. The deflecting magnet is 4 inches long; it is a hollow cylinder, carrying in its internal tube a collimator, by means of which its time of vibration is observed in another apparatus. The frame which supports the suspension-piece of the deflected magnet carries also the telescope directed to the magnet-mirror; it rotates round the vertical axis of a horizontal graduated circle whose external diameter is 10 inches. The deflecting magnet is always placed on the E. or W. side of the deflected magnet, with one end towards the deflected magnet. In the reduction of the observations, the precepts contained in the Skeleton Form prepared by the Kew Observatory have received the strictest attention.

The following is the explanation of the method of reduction.

The distance of the centers of the deflected and deflecting magnet being known, it is supposed (from observations made at Kew, of which the details have not reached me) that the magnetism of the deflecting magnet is so altered by induction that the following multipliers ought to be used in computing the Absolute Force $:$ --

The correction of the magnetic power for temperature t_0 of Fahrenheit, reducing all to 35° of Fahrenheit, is

 o '000131261 (t_0 -35) + 0 '000000259 (t_0 -35)²

 A_1 is $\frac{1}{2}$ (distance)⁸ × sine deflection, corrected by the two last-mentioned quantities, for distance 1 foot; A_2 is the similar expression for distance 1.3 foot; A'_2 is $\frac{A_2}{(1\cdot 3)^2}$; P is $\frac{A_1 - A_2}{A_1 - A_2}$. A mean value of P is adopted from various observations; then $\frac{m}{\overline{X}} = A_1 \times \left(1 - \frac{P}{1}\right)$ for smaller distance, or $= A_2 \times \left(1 - \frac{P}{1.69}\right)$ for larger distance. The mean of these is usually adopted for the true value of $\frac{m}{X}$.

. ABSOLUTE MEASURE OF HORIZONTAL MAGNETIC FORCE: TABLES OF REDUCTIONS OF THE MAGNETIC OBSERVATIONS. *xx.xix*

For computing the value of mX from observed vibrations, it is necessary to know K , the moment of inertia of the magnet as mounted. The value of log. $\pi^2 K$ furnished by Mr. Stewart is 1.66073 at temperature 30° and 1.66109 at temperature 90°. Then putting T for the time of the magnet's vibration as corrected for induction, temperature, and torsion-force, the value of mX is = $\frac{\pi^2 K}{T^2}$. From the combination of this value of mX with the former value of $\frac{m}{\overline{X}}, m$ and X are immediately found.

It appears, from a comparison of observations given in the Introduction to the *Magnetical and Meteorological Observations,* 1862, that the determinations with the Old Instrument (in use to 1861) ought to be diminished by $\frac{1}{117}$ part, to make them comparable with those of the Kew Unifilar.

The computation of the values of *m* and X has, to the year 1857, been made in reference to English measure only, using the foot and the grain as the units of length and weight; but, for comparison with foreign observations of the Absolute Intensity of Magnetism, it is desirable that *X* should be expressed also in reference to Metric measure, in terms of the millimetre and milligramme. If an English foot be supposed equal to α times the millimetre, and a grain be equal to β times the milligramme, then it is seen that, for the reduction of $\frac{m}{X}$ and mX to Metric measure, these must be multiplied by α^3 and $\alpha^2\beta$ respectively. Hence X^2 must be multiplied by $\frac{\beta}{\alpha}$, and X by $\sqrt{\frac{\beta}{a}}$. Assuming that the mètre is equal to 39.37079 inches, and the gramme equal to 15'43249 grains, log. $\sqrt{\frac{\beta}{\alpha}}$ will be found to be = 9'6637805, and the factor for reducing the English values of X to Metric values will be 0.46108 or $\frac{1}{2.1689}$. The values of X in Metric measure thus derived from those in English measure are given in the proper table,

:§ 11. *Explanation of the Tables Q/ Reductions* Q/ *ihe Magnetic Observations (excluding the days of great Magnetic Disturbance).*

 \mathbb{S}^1 . All \mathcal{M}^1 is an

 $\gamma_{\rm{max}}$

The Indications, on which the reductions of this section and the next are founded, are derived entirely from the measures of the ordinates of the Photographic Curves.

The first step taken was to divide the days of observation into two groups; in one of which the magnetism was generally so tranquil that it appeared proper to use those days for determination of the laws of diurnal inequality; while in the other group the movements of the magnetic instruments were so violent, and the photographic curves . traced' by them so irregular, that it appeared impossible to employ them, except by the exhibition of every motion of the magnet during the day. A similar division into groups had been made in two Memoirs printed in the Philosophical Transactions.

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For the year 1870, the following days, sixteen in number, were selected by Mr. Glaisher as exhibiting practically the same amount of irregularity which he had considered as defining the class of Days of Great Disturbance in the Memoirs to which I have alluded:-

February 1, 11, March 21, April 5, May 20, August 19, 20, September 3, 24, 25, 26, October 14, 24, 25, November 8, December 17.

These days being separated, the photographic sheets for the remaining days were thus treated. Through each photographic curve a pencil line was drawn, representing, as well as could be judged, the general form of the curve without its petty irregularities. These pencil curves only were then used; and their ordinates were measured, with the proper pasteboard scales, at every hour. The methods of forming from these the various tables of this section require no special explanation.

The temperature of the Magnetometers was maintained in so great uniformity through each day that no apprehension is entertained of the slightest appreciable error in the diurnal inequalities of horizontal force and vertical force, as a consequence of the omission of temperature-correction. But it was impossible to maintain perfect uniformity of temperature through all the seasons. I have, therefore, exhibited, in the Tables of Mean Force in each month, the mean temperature of the month. It will be borne in mind, therefore, that the numbers exhibited are *not* corrected for temperature, but require the correction corresponding to the printed mean temperatures.

§ 12. *Explanation* of. *tIle Tables* of *Indications* if *Magnetometers on sixteen days cif Great Magnetic Disturbance.*

Telescope-observations of the Magnetometers have usually been made four times every day, except on Sundays, on which days two or three observations only have been taken; but, though these observations are employed in forming the base lines on the photographic sheets, their immediate results are not necessarily given in the Tables.

For each photographic record, a new base-line, representing a convenient reading in round numbers of the element to which it applies, has been drawn on the sheet. Then the Assistant, who is charged with the translation of the curve-ordinates into numbers, remarks the salient points of the curve, or the points which if connected by straight lines would produce a polygon not sensibly differing from the photographic curve; to each of these he applies the scale of pasteboard or glass proper for the element under consideration; the base of the scale determines the time on the timescale, and the reading of the scale for the point of the photographic curve gives the quantity which is to be added to the value for the new base-line. The ordinate-

TABLES OF INDICATIONS OF THE MAGNETOMETERS: REGISTER OF SPONTANEOUS TERRESTRIAL GALVANIC CURRENTS. xli

reading so formed is printed without alteration in the Tables. It is particularly to be remarked that the indications for horizontal force and vertical force are *not corrected for temperature.*

In preceding years, allusion has been made to the occasional dislocations of the curve of Vertical Force. No instance of such dislocation has presented itself in 1870. It is believed that these dislocations were produced by bringing a magnet into the proximity (though not very close) of the magnetometer; and this supposed cause of error has, in late years, been carefully avoided.

§ 13. *Wires and Photographic self-registering Apparatus for continuous Record of Spontaneous Terrestrial Galvanic Currents.*

In order to obtain an exhibition of the spontaneous galvanic currents which in some measure are almost always discoverable in the earth, and which occasionally are very powerful, it was necessary to extend two insulated wires from an earth connexion at the Royal Observatory, in two directions nearly at right angles to each other, to considerable distances, where they would again make connexion with the earth. By the kindness of the Directors of the South Eastern Railway Company, to whom the Royal Observatory has on several occasions been deeply indebted, two connexions were made; one to a station near Dartford, at the direct distance $9\frac{3}{1}$ miles nearly, in azimuth (measured from North, to East, South, West), 102° astronomical or 122° magnetical, the length of the connecting wire being about $15\frac{3}{5}$ miles; the other to a station near Croydon, at the direct distance 8 miles, in azimuth, 209° astronomical, or 229° magnetical, the length of the connecting wire being about $10\frac{1}{2}$ miles. At these two stations connexion was made with earth. The details of the course were as follows. The wires were soldered to a water pipe in the Magnetic Ground at the Royal Observatory. Thence they entered the Magnetic Basement, and passed through the photographic selfregistering apparatus (to be shortly described). From it they were led up the electrometer mast to a height exceeding 50 feet, and thence they were swung across the grounds to a chimney above the Octagon Room. They descended thence, and were led to a terminal board in the Astronomical Computing Room, to which an intermediate galvanometer can be attached for eye-observation of the currents. From this point they were led to the "Battery Basement," and, with other wires, passed under the Park to the Greenwich Railway Station, and upon the telegraph poles. One wire branched off at the junction with the North Kent Railway to Dartford, the other at the junction with the Croydon Branch Railway to Croydon. At both places their connexion with earth was made by soldering to water-pipes, as at the Royal Observatory.

These wires remained in the places described till the end of 1867. It had been discovered in experience that a much smaller separation of the extreme points of earth-connexion would suffice, and it was conjectured that advantage might arise from making the two earth-connexions of each wire on opposite sides of the Observatory and nearly equidistant from it, instead of making one earth-connexion of each within

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$xlii$ INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

the Observatory grounds. In 1868, therefore, the following wire-courses were substituted. One wire is connected with earth, by a copper plate, at the Lady Well station of the Mid-Kent Railway; it is thence led by a circuitous course to the North Kent Junction with the Greenwich Railway, to the Royal Observatory (for communication with the self-registering apparatus), back to the North Kent Junction, then by North Kent Railway and Angerstein Branch to the Angerstein Wharf, where it is connected with earth by a copper plate. The other wire is connected with earth by a copper plate at the North Kent Junction, then passes to the Royal Observatory and back to the Junction, and then along the North Kent Railway to the Morden College end of the Blackheath Tunnel, where it is connected with earth in the same manner. The straight lines connecting the extreme points of the wires cross each other near the middle of their lengths and near the Royal Observatory; the length of the first line is nearly 3 miles, and its azimuth 56° N. to E. (magnetic); that of the second line is nearly $2\frac{1}{2}$ miles, and its azimuth 136°. But, in the circuitous courses above described, the length of the first wire is about $10\frac{3}{8}$ miles, and that of the second $6\frac{1}{4}$ miles. These wires were established and brought into use on 1868, August 20.

The apparatus for receiving the effects of the galvanic currents consists essentially of two magnetic needles (one for each wire), each suspended by a hair so as to vibrate horizontally within a galvanic coil, exactly as in the ordinary speaking telegraph (supposed to be laid horizontally) ; these coils being respectively in the courses of the two long wires. A current of one kind, in either wire, causes the corresponding needle to turn itself through an angle nearly proportioned to the strength of the current, in one direction; a current of the opposite kind causes it to turn in the opposite direction. These turnings are registered by the following apparatus.

The carrier of each magnet carries also a small plane mirror, which receives all the azimuthal motions of the magnet. The light of a gas-lamp passes, through a minute aperture, and shines upon the mirror; the divergent pencil is converted into a convergent pencil by refraction through crossed cylindrical lenses (with axes vertical before the pencil reaches the mirror, and with axes horizontal where the pencil is received from the mirror), which, under the circumstances, were more convenient than spherical lenses. A spot of light is thus formed upon the photographic paper wrapped upon a cylinder of ebonite, which is covered by a glass cylinder, and made to rotate in twenty-four hours by clock-work, exactly as for the register of the magnetic elements. As in the case of declination and horizontal-force, the two earth currents make their registers upon opposite sides of the same barrel, and upon different parts of the sheet; the same gaslight serving for the illumination of both.

A portion of a base-line for either record is obtained at any time by simply breaking the galvanic communication.

The photograph records were regularly made, with the wires in the first position, from 1865, March 15, to the end of 1867. Fifty-three days, on which the magnetic disturbances were active, were selected for special examination; and for these the equivalent galvanic cnrrents in the north and wcst directions were computed, and their

ApPARATUS FOR SPONTANEOUS TERRESTRIAL GALVANIC OS FOR SPONTANEOUS LERRESTRIAL GALVANIC $xliii$
CURRENTS: STANDARD BAROMETER.

effects in producing apparent magnetic disturbances in the west and north directions were inferred. They correspond almost exactly with those indicated by the magnetometers. Then the records for all the days of tranquil magnetism were reduced in the same manner, not for comparison with the magnetometer-results, but for ascertaining the diurnal laws of the galvanic currents. These laws were found to be very different from the laws of magnetic diurnal inequalities. These discussions have been communicated to the Royal Society in two papers, of which the first is printed in the Philosophical Transactions, 1868.

The records with the wires in the new positions have been regularly made since 1868, August 20, but have not yet been discussed.

§ 14. *Standard Barometer.*

The Barometer is a standard, by Newman, mouuted in 1840. It is fixed on the South wall of the West arm of the Magnetic Observatory. The graduated scale which measures the height of the mercury is made of brass, and to it is affixed a brass rod, passing down the inside of one of the upright supports, and terminating in a . conical point of ivory; this point in observation is made just to touch the surface of the mercury in the cistern, and the contact is easily seen by the reflected and the actual point appearing *Just* to meet each other. The rod and scale are made to slide up and down by means of a slow-motion screw. The scale is divided to 0^{in} 05.

The vernier subdivides the scale divisions to $0^{\text{in}}002$; it is moved by a slow-motion screw, and in observation is adjusted so that the ray of light, passing under the back and front of the semi-cylindrical plate carried by the vernier, is a tangent to the highest part of the convex surface of the mercury in the tube.

The tube is $0^{\text{in}} 565$ in diameter; the correction for the effect of capillary attraction is therefore only $+ 0^{in}$ 002. The cistern is of glass.

At the bottom of the instrument are three screws, turning in the fixed part of the support, and acting on the piece in which the lower pivot of the barometer-frame turns, for adjustment to verticality: this adjustment is examined weekly.

The readings of this barometer, until 1866, August 20^d , 0^h , are considered to be coincident with those of the Royal Society's flint-glass standard barometer. On that day a change was made in the barometer. It had been remarked that the slow-motionscrew at the bottom of the sliding rod (for adjusting the ivory point to the surface of the mercury in the cistern) was partly worn away: and on August 20 the sliding rod was removed from the barometer by Mr. Zambra to remedy this defect. It was restored on 1866, August 30^d , 3^h . Before the removal of the sliding rod, barometric comparisons had been made with a standard barometer the property of Messrs. Murray and Heath, and with two barometers, Negretti and Zambra, Nos. 646 and 647. While the sliding rod of the Greenwich standard was removed, Negretti and Zambra 647 was used for daily observations. After the new equipment of the standard barometer,

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another series of comparisons with the same barometers was made: from which it was found (the three auxiliaries giving accordant results) that the readings of the barometer, in its new state, required a correction of $-0ⁱⁿ \cdot 006$. This is applied in the printed observations commencing with 1866, August 30.

The height of the cistern above the mean level of the sea is 159 feet. This element is founded upon the determination of Mr. Lloyd, in the *Phil. Trans.,* 1331 ; theelevation of the cistern above the brass piece inserted in a stone in the transit-room (to which Mr. Lloyd refers) being $5^{\text{ft}}.2^{\text{in}}$.

The barometer has been read at $21^{\rm h}$, $0^{\rm h}$, $3^{\rm h}$, $9^{\rm h}$ (astronomical), on every day, excepting on Sundays, and on Good Friday and Christmas Day, on which days fewer observations have been taken. Every reading has been reduced to the reading which would have been obtained at the temperature 32° of the mercury and scale, by application of the correction given in Table IL (pages 82 to 87) of the Report of the Committee of Physics of the Royal Society. The mean of the reduced readings has then been taken for eaeh civil day, and finally converted into mean daily reading, by application of the correction inferred from Mr. Glaisher's paper in the *Philosophical Transactions*, 1848, Part I, Table I, page 127.

In the printed record of the barometrical and all other meteorological observations, the day is to be understood, generally, as defined in civil reckoning.

§ 15. *Photo{{raphic self-registering Apparatus for continuous Record qf the Readings* of *the Barometer.*

The Photographic self-registering Apparatus for continuous Record of Magnetic Vertical Force is furnished (as has been stated) with a vertical cylinder covered with photographic paper and revolving in 24 hours. North of the surface of this cylinder, at the distance of about 30 inches, is a large syphon barometer, the bore of the upper and lower extremities of its arms being about 1'1 inch. A glass float partly immersed in the quicksilver of the lower extremity is partially supported by a counterpoise acting on a light lever (which turns on delicate pivots), so that the wire supporting the float is constantly stretched, leaving a definite part of the weight of the float to be supported by the quicksilver. This lever is lengthened to carry a vertical plate of opaque mica with a small aperture, whose distance from the fulcrum is nearly eight times the distance of the point of attachment of the float wire, and whose movement, therefore, is nearly four times the movement of the column of a cistern-barometer. Through this hole the light of a lamp, collected by a cylindrical lens, shines upon the photographic paper.

The scale of time is established by means of occasional interruptions of the light, and the scale of measnre is established by comparison with occasional eye-observations.

This barometer was brought into use in 1848, but its indications were not satis-

PHOTOGRAPHIC BAROMETER; DRY-BULB AND WET-BULB THERMOMETERS. *xlv*

factory till the mercury was boiled in the tube by Messrs. Negretti and Zambra on 1853, August 18, since which time they have appeared unexceptionable. Results of the indications are printed in the *Maxima and Minima of the Barometer*, near the end of the Meteorological Results.

§ 16. *Thermometers for ordinary Ohservation* cif *the Temperature* if *the Air and Evaporation.*

The Dry-Bulb Thermometer, the Wet-Bulb Thermometer, the Maximum Self-Registering Thermometers, both dry and wet, and the Minimum Self-Registering Thermometers, dry and wet, all for determination of the temperature of the air and of evaporation, are mounted on a revolving frame whose fixed vertical axis is planted in the ground. From the year 1846 to 1863 the post forming the vertical axis was about 23 feet south (magnetic) of the S.S.E. angle of the south arm of the Magnetic Observatory; in 1863 it was moved to a position about 35 feet south (astronomical) of the south angle. A frame revolves on this post, consisting of a horizontal board as base, of a vertical board projecting upwards from it connected with one edge of the horizontal board, and of two parallel inclined boards (separated about three inches) connected at the top with the vertical board, and at the bottom with the other edge of the horizontal board. The outer inclined board is covered with zinc. The air passes freely between all these boards.

The dry and wet-bulb thermometers are attached to the outside, and near the center of the vertical board; the maximum and minimum thermometers for air towards one vertical edge, and those for evaporation towards the other vertical edge, with their bulbs at almost the same level, and near to those of the dry and wet-bulb thermometers; their bulbs are about 4 feet above the ground and projecting from 2 inches to 3 inches below the horizontal board. Above the thermometers is a small projecting roof to protect thern from rain. The frame is always turned with the inclined side towards the sun. It is presumed that the thermometers are thus sufficiently protected.

The graduations of all the thermometers used in the Royal Observatory rest fundamentally upon those of a Standard Thermometer, the property of Mr. Glaisher, which derives its authority from comparison with original thermometers constructed by the late Rev. R. Sheepshanks about the years 1840-1843, in the course of his preparations for the construction of the National Standard of Length. The whole of the radical determinations of Freezing Point, Boiling Point, and Subdivision of Volume of Tube, were made by Mr. Sheepshanks with the utmost care: it is believed that these were the first original thermometers that had been constructed in England for many years. Mr. Glaisher's thermometer has been adopted as the standard of reference for all the thermometers used in the Royal Observatory since 1840.

The Dry-Bulb Thermometer *is* by Newman. 'fhe corrections required for its

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readings, as found by comparison with the standard above-mentioned, are as $follows :=$

The wet-bulb thermometer is by Negretti and Zambra, and is in every respect similar to the dry-bulb thermometer. The corrections required to the readings of this thermometer are--

Dry-bulb and wet-bulb thermometers, with pea-bulbs and porcelain scales, Negretti and Zambra 1179, are also mounted on the roof of. the library, 4 feet above the leads and 22 feet above the ground. No corrections for index error are applied to the readings of these thermometers.

On 1869, September 30, dry-bulb and wet-bulb thermometers were mounted on the roof of the cabinet containing the registering mechanism of Robinson's Anemometer, but below the revolving cups, at the height 4 feet above the flat roof and 50 feet above the ground. No corrections for index errors are applied to their readings.

1'he eye-readings of the dry-bulb and wet-bulb thermometers have usually been taken at the hours (astronomical reckoning) $21^{\rm h}$, $0^{\rm h}$, $3^{\rm h}$, $9^{\rm h}$, and corrected by application of the numbers given above. They are not printed in the present volume.

The dew-point has been inferred exclusively from the simultaneous observations of the dry-bulb and wet-bulb thermometers, by multiplying the difference between the readings of these thermometers by a factor peculiar to the temperature of the air, and subtracting the product from the reading of the dry-bulb thermometer.

DRY-BULB AND WET-BULB THERMOMETERS: DEW POINT. xlvii

These factors have been found by Mr. Glaisher from the comparison of a great number of dew-point determinations, obtained by use of Daniell's hygrometer, with simultaneous observations of dry-bulb and wet-bulb thermometers. The first part of this investigation was published in full, in the volume of *Magnetical and Meteorological Observations* for 1844, pages 67-72; it was based upon all the observations made up to that time. Subsequently, the comparison was extended to include all the simultaneous observations of these instruments made at the Royal Observatory, Greenwich, from 1841 to 1854, with some observations taken at high temperatures in India, and others at low and medium temperatures at Toronto. The results at the same temperature were found to be the same at these different localities, so far as the climatic circumstances permitted comparison. (See Glaisher's Hygrometrical Tables, 5th Edition). The following table exhibits the result of the entire comparison; it has been used in forming the dew-points in the present volume.

TABLE OF FACTORS by which the DIFFERENCE of READINGS of the DRy-BuLB and WET-BuLB THER-MOMETERS is to be MULTIPLIED in order to PRODUCE the DIFFERENCE between the READINGS of the DRy-BuLB and DEW-POINT THERMOMETERS.

Reading of Dry-bulb Thermometer.	Factor.	Reading of Dry-bulb Thermometer.	Factor.	Reading of Dry-bulb Thermometer.	Factor.	Reading of Dry-bulb Thermometer.	Factor.
\mathbf{o} 10 11 12 13 14 15 16 17 18 19 20 2I 22 -23 24 $2\overline{5}$ 26 27 28 29 30 31 32	8.78 8.78 8.78 8.77 8.76 8.75 8 70 8.5 62 8 5 _o 34 8 8. 14 88 7° 60 7. 28 6. 92 6.53 \circ 8 6۰ 5.61 5.12 4.63 4.15 3.70 3.32	3 ^o 34 35 36 37 38 39 40 41 42 4 ³ 44 4 ₅ 46 47 48 49 5 _o 51 52 53 54 55	3.01 2.77 2.60 2 ° 50 2.42 2.36 2.32 29 \mathbf{z} 2.26 2.23 2.20 2.18 2.16 2^{\cdot} 14 2.12 2.10 2.08 2.06 2.04 2.02 2.00 1.98 1. ⁹⁶	$5\overset{\circ}{6}$ 57 58 59 60 61 62 63 64 65 66 67 68 69 70 71 72 7 ³ 74 75 76 77 78	I'94 1.92 1.90 8۰. 1 1.88 \cdot 87 r 1.86 1.85 1.83 1.82 1.81 1.80 1.79 1.78 1'77 1.76 1.75 1.74 1.73 1.72 1'71 1.70 1.69	۰ 79 80 81 82 83 84 85 86 87 88 89 90 91 92 9 ³ 94 $95 -$ 96 97 98 99 100	1.69 1.68 ı 68 1.67 1.67 1:66 1.65 1.65 1.64 1.64 1.63 1.63 1.62 1.62 1.61 1.60 60۰ ı .59 1 1.59 1.58 1.58 1.57

The maximum self-registering thermometer is a mercurial thermometer, of the construction invented by Messrs. Negretti and Zambra. There is a small detached piece of glass in the tube, just above a bent part of the tube (near the bulb), through which the piece of glass cannot pass down. The column of mercury in rising lifts the glass' up and passes freely; but in descending it is unable to pass the glass, and the lower mass of mercury descends, leaving a vacant space below the glass, and

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leaving a portion of the mercury above it. The piece of glass operates as an efficient eaving a portion of the mercury above it. The piece of glass operates as an valve. The corrections to the readings of this thermometer are as follows :-000

There is a similar thermometer for the maximum wet-bulb reading (Negretti and Zambra No. 7537): no corrections have been applied to its readings.

The minimum self-registering thermometers are alcohol thermometers, of the construction known as Rutherford's. Λ sliding glass index allows the alcohol in rising to pass above it, but is drawn down by the peculiar action of the bounding surface of the fluid when it sinks. The readings of that which gives the minimum temperature of the air require no correction.

The minimum wet-bulb thermometer (Negretti and Zambra, No. 3627) is also free from sensible error.

The mean daily values of dry thermometer in the printed columns are found by combining two results derived from different sources. The first and simpler result is the mean of the maximum and minimum, corrected by a small quantity depending on the month, given in Table III. of Mr. Glaisher's paper in the Philosophical *Transactions,* 1848, page 130. The second result is formed by taking the means of the four eye-observations at 21^{h} , 0^{h} , 3^{h} , 9^{h} , and applying a correction thus investigated. The daily range being found by taking the difference between the maximum and minimum, this daily range is multiplied by the mean of the factors in Table IV. of Mr. Glaisher's paper before mentioned corresponding to the hours of observation; the application of this correction to the mean of the eye-observations gives the second result. (It is evident that this process is applicable to any number of eyeobservations.) These two results are then combined to form a mean, weights being given proportional to the number of observations contributing to each result.

For the mean daily value of dew point, the usual process is,—by observing the difference between dry and wet thermometers, and by use of the table of factors printed in page *xlvii* above, to form the difference between air-temperature and dew point at each of the hours of reading; to take the mean of the deduced dew-points; and to apply a correction which is the mean of the corrections in Mr. Glaisher's Table VIII. for the scverai hours of observation. Sometimes, however, the following process is used. 'fhe correction for diurnal range applicable to the lnean of the eye-observations of the dry thermometer having been found (as is described above), this correction is multiplied by a fraction, whose numerator is the mean of corrections to wet bulb thermometer in Table VII. for the hours of observations, and whose denominator is the mean of corrections to dry thermometer in Table II. for the same hours; and thus a correction is found which is applied to the mean of the eye-observations of wet bulb

MAXIMUM AND MINIMUM THERMOMETERS: MEAN DAILY VALUES OF DRY THERMOMETER AND DEW.POINT: *Xli\$* PHOTOGRAPHIC THERMOMETERS.

thermometer, to form the mean wet bulb for the day. Then by use of the mean dry bulb reading for the day and the mean wet bulb reading for the day and the table of factors above, the mean dew point for the day is formed.

§ 17. *Photographio self-registering Apparatus for oontinuous Reoord* of *the Readings cifthe Dry-Bulb and Wet-Bulb Thermometers.*

About 28 feet south (magnetic) of the south-east angle of the south arm of the Magnetic Observatory, and about 25 feet east of the thermometers for eye-observations, is a shed lOft. 6 in. square, standing upon posts 8 feet high, under which are placed the photographic thermometers, the dry-bulb thermometer towards the east, and the wet-bulb thermometer towards the west. The bulbs of the thermometers are 8 inches in length, and 0·4 inch internal bore, and their centers are about 4 feet above the ground. The bulb of one of the thermometers is covered with tnuslin throughout its whole length, which is kept moist by means of capillary passage of water along cotton wicks leading to a vessel filled with water.

There are small adjustments admitting the raising or dropping of the thermometers, so that the register of their changing readings may be on a convenient part of the paper. The thermometer frames are covered by plates having longitudinal apertures, so narrow, that any light which may pass through them is completely, or almost completely, intercepted by the broad flat column of mercury in the thermometer-tube. Across these plates a fine wire is placed at every degree; and at the decades of the degrees, and also at 32°, 52°, and 72°, a coarser wire is placed. A gas lamp is placed about 9 inches from each thermometer (east of the dry bulb and west of the wet bulb), and its light, condensed by a cylindrical lens, whose axis is vertical, shines through the thermometer-tube above the surface of the mercury, and forms a well-defined line of light upon the photographic paper, which is wrapped around the cylinder. The axis of this cylinder is vertical; its mounting is in all respects similar to that of the Vertical Force cylinder. As the cylinder, covered with photographic paper, revolves under the light, which passes through the thermometer-tube, it receives a broad sheet of photographic trace, whose breadth (in the direction of the axis of the cylinder) varies with the varying height of the mercury in the thermometer-tube. The light in its passage is intercepted by the wires placed across the tube at every degree, and there are, there. fore, left upon the paper corresponding lines in which there is no photogenic action.

The cylinder revolves in 48 hours; the daily photographic traces of the two thermometers are thus simultaneously registered on opposite sides of the cylinder without intermixing. The length of the glass cylinder used till 1869, March, is $13\frac{1}{9}$ inches, and its circumference is about 19 inches. On 1869, March 5, an ebonite cylinder was

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introduced, whose length is 10 inches, and circumference about 19 inches; and at a later time the cylinder was made to revolve in 50 hours instead of 48 hours, to insure the separation of the records of the two thermometers. **是然后, No. 1000 (1)**

§ 18. *Thermometers for Solar Radiation and Radiation to the Sky.*

The thermometer for Solar Radiation, which to the end of the year 1864 was placed in an open box about 10 feet south of the south-west angle of the south arm of the Magnetic Observatory, is now laid on the grass, near the same place.

The thermometer is a self-registering maximum mercurial thermometer of Negretti and Zambra's construction; its bulb is blackened, and enclosed in a glass sphere from which the air has been exhausted. Its graduations are correct, and the numbers inserted in the tables are those read from the instrument without alteration. The thermometer is read at 9^{h} a.m., noon, 3^{h} p.m., and occasionally at 9^{h} p.m.; the highest of these readings is adopted as the maximum for the day.

The use of a thermometer with blackened bulb not inclosed in an exhausted sphere was discontinued at the end of 1865.

The thermometer for radiation to the sky is placed near to the Solar Radiation thermometer, with its bulb resting on short grass, and fully exposed to the sky. It is a self-registering minimum spirit thermometer of Rutherford's construction, made by Negretti and Zambra. Its graduation is correct, and the numbers inserted in the table are those read from the scale without alteration. It is read every day at 9^{h} a.m., and occasionally at 9^h p.m.

§ 19. *Thermometers sunk below the Surface cifthe Soil at'different Deyths.*

These thermometers were made by Messrs. Adie of Edinburgh, under the immediate superintendence of the late Professor J. D. Forbes. The graduation was made by Professor Forbes himself.

The thermometers are four in number. They are all placed in one hole in the ground, the diameter of which in its upper half is 1 foot, and in its lower half about 6 inches. Each thermometer is attached in its whole length to a slender piece of wood, which is planted in the hole with it. The place of the hole is 20 feet south of the extremity of the south arm of the Magnetic Observatory, and opposite the center of its south front.

The soil consisted of beds of sand; of flint-gravel with a large proportion of sand; and of flints with a small proportion of sand, cemented almost to the consistency of

RADIATION THERMOMETERS: DEEP-SUNK THERMOMETERS.

pudding-stone. Every part of the gravel and sand extracted from the hole was perfectly dry.

The bulbs of the thermometers are cylindrical, 10 or 12 inches long and 2 or 3 inches in diameter. The bore of the principal part of the tubes, from the bulb to the graduated scale, is very small. In that part to which the scale is attached, the tube is larger.

. The thermometer No. 1 was dropped into the hole to such a depth that the center of its bulb was 24 French feet (25'6 English feet) below the surface: then dry sand was poured in till the hole was filled to nearly half its height. Then No. 2 was dropped in till the center of its bulb was 12 French feet below the surface; No. 3 and No. 4 till the centers of their bulbs were respectively 6 and 3 French feet below the surface; and the hole was then completely filled with dry sand. The upper parts of the tubes, carrying the scales, were left projecting above the surface: No.1 by 27·5 inches, No.2 by 28·0 inches, No.3 by 30·0 inches, and No.4 by 32'0 inches. Of these lengths, the parts 8.5 , 10.0 , 11.0 , and 14.5 inches, respectively are tube with narrow bore.

The projecting parts of the tubes are protected by a wooden case or box fixed to the ground; the sides of the box are perforated with numerous holes, and it has a double roof. In the North face of this box is a large plate of glass through which the thermometers are read. Within the box are two smaller thermometers, one $(No.5)$ whose bulb is sunk one inch in the ground, and one $(No.6)$ whose bulb is in the free air nearly in the center of the box.

The fluid of the four long thermometers is alcohol tinged with a red colour.

The lengths of 1° on the scales of Nos. 1, 2, 3 and 4, are respectively 2^{in} , 1^{in} , 0^{in} , 9 , and $0^{\text{in}}.55$; and the ranges of the scales, as first mounted, were, 43° 0 to 52° 7 , 42° 0 to *56°'8,* 39°'0 to *57°'5,* and *34°·2* to 64°·5.

These ranges for Nos. 2, 3, and 4, were found to be insufficient in some years, particularly those of Nos, 3, and 4, or the thermometers sunk to the depth of 6 feet and 3 feet.

In 1857, June 22, Messrs. Negretti and Zambra removed from Nos. 3 and 4 a quantity of fluid corresponding to the extent of 5° on their scales, and the scales of these two thermometers were then lowered by that linear extent, making the readings the same as before. Their ranges are now, respectively, 44° to 62° -5, and 39° -2 to 69° -5.

In subsequent years it was found that the amount of fluid removed was somewhat too great, for now at the lower end of the scale the 6-foot thermometer sometimes falls below the limit of its scale or 44° ; and the 3-foot thermometer below $39^{\circ}\cdot 0$; in which cases the alcohol sinks into the capillary tube.

The readings at the early part of the series were at times defective at high temperatures, but always complete at low temperatures; now, they are generally complete at

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high temperatures, and are at times defective at low temperatures. The two combined, however, will enable us to complete all readings.

These thermometers are read once a day, at noon, and the readings appear in the printed volumes as read from their scales without correction.

On 1869, July 21, Mr. Zambra removed fluid from No. 1 to the amount of 2° . and from No. 2 to the amount of 1° -5, and inserted in No. 4 fluid to the amount of 1^c 5. The scales were re-engraved, to make the reading at every temperature the same as before.

§ 20. *Thermometers immersed in the Water* if *the Thames.*

The self-registering maximum and minimum thermometers for determining the highest and lowest temperatures of the water of the Thames are by Messrs. Negretti and Zambra, and are observed every day at 9^h a.m.

A strong wooden trunk is firmly fixed to the side of the Dreadnought Hospital Ship, about 5 feet in length, and closed at the bottom; the bottom and the sides, to the height of 3 feet, are perforated with a great number of holes, so that the water can easily flow through; the thermometers are suspended within this trunk *eo* as to be about 2 feet below the surface of the water, and I foot from the bottom of the trunk.

The regular observations were made under the superintendence of the Medical Officers of the Ship till April 12. On April 13 the thermometers were withdrawn, in consequence of the Dreadnought ceasing to be used as a hospital ship; and from that day to the end of the year no observations were made.

No readings were taken on January 2, February 19, 26, March 6, 7, 21, and April 7.

The index-error corrections to these thermometers were :—

$§ 21.$ *Osler's Anemometer.*

This anemometer is self-registering: it was made by Newman, on a plan furnished by A. Follett Osler, Esq., F.R.S., but has received several changes since it was originally constructed. A large vane, which is turned by the wind, and from which a vertical spindle proceeds down nearly to the table in the north-western turret of the ancient part of the Observatory, gives motion by a pinion upon the spindle to a rackwork carrying a pencil. This pencil makes a mark upon a paper affixed to a board

THAMES THERMOMETERS: OSLER'S ANEMOMETER. *liii*

which is moved uniformly in a direction transverse to the direction of the rack-motion. The movement of the board is effected by means of a second rack connected with the pinion of a clock. The paper has lines printed upon it corresponding to the positions which the pencil must take when the direction of the vane is N , E , S , or W , and also has transversal lines corresponding to the positions of the pencil at every hour. The first adjustment for azimuth was obtained by observing from a certain point the time of passage of a star behind the vane-shaft, and computing from that observation the azimuth; then on a calm day drawing the vane by a cord to that position, and adjusting the rack, &c., so that the pencil position on the sheet corresponded to that azimuth.

This construction originally arranged by Mr. Osler was in use till the middle of 1866, when the following modifications were made in it by Mr. Browning:—

The vane-shaft was made to bear upon anti-friction-rollers running in a cup of oil. For elucidation of the following description of the apparatus which it carries, I refer to Figure 3 on the engraving at the end of the Introduction to the volume of 1866. To the vane-shaft is attached!a rectangular frame C, which rotates with the vane. To this frame are firmly attached the ends of four strong springs D, which rise from the point of attachment in a vertical direction, are then bent so as to descend below the frame C , and are then bent upwards so as to rise a short distance, where they terminate, each of them thus fonning a large hook. 'fo the interior of each strong spring, near to its upper bend, is affixed a very weak spring, which descends free into the lower bend or hook of the strong spring, so that its lower end may be rnoved by a light pressure till it reaches and takes bearing against the bent-up part of the strong spring, after which it cannot be further moved without moving the strong spring, and will therefore require much greater pressure. . The four ends of these four light springs carry the circular pressureplate A by the following connexions. The two which are farthest from A , or which are below the wide part of the vane, are united by a light horizontal cross-bar G; and from the ends of these springs proceed four light bars E, which are attached to points of the pressure-p1ate A, near its circumference. The two. ends of light springs which are nearest to A are also united by a light horizontal cross bar, which is attached to a projection from the center of the plate A. (The diagonal lines upon A, in the diagram, represent indistinctly two strengthening edge-bars upon the pressure-plate, and the projection above-mentioned is fixed to their intersection.) The weight of the pressure-plate thus rests entirely on the slender springs; it is held steadily in position, as regards the opposition to the wind, and it moves without sensible friction. A light wind drives it through a considerable space, until the ends of one pair of light springs touch their large hooks; then for every additional pound of pressure the movement is smaller, till the ends of the other pair of light springs touch their large hooks; after this the movement for every additional pound of pressure is still further diminished. This apparatus was arranged by Mr. Browning. The communication with the pencil below is similar

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to that in the first construction: the cord and pulley are omitted in the drawing to avoid confusion.

The pressure-pencil below is carried by a radial bar, whose length is parallel to the scale of hours; it is brought to zero by a small weight on a cord running over a pulley.

The surface of the pressure-plate is 2 square feet, or double that in the old construction. The scale of indications on the recording-sheet was determined experimentally as in the old instrument; yet it is remarked that the pressures of wind per square foot appear generally greater than formerly.

The scale for small pressures is much larger, and their indications much more certain than formerly. A pressure of an ounce per square foot is clearly shown.

A rain gauge of peculiar construction is carried by this instrument, by which the fall of rain is registered with reference to the time of the fall. It is described in \S 23.

A fresh sheet of paper is applied to this instrument every day at 22^h mean solar time.

§ 22. *Robin.r;on's Anemometer.*

In the latter part of the year 1866, a new instrument, on the principles described by Dr. Robinson in the Transactions of the Royal Irish Academy, vol. xxii., adapted to give a continuous record of the velocity of the wind, was mounted by Mr. Browning, of which the principal parts are represented in Figures 1 and 2 of the engraving in the Introduction 1866. The motion is given (as in the former instrument) by the pressure of the air on four hemispherical cups, the distance of the center of each from the axis of rotation being 15·00 inches. The foot of the axis is a hollow flat one bearing upon a sharp cone which rises up from the base of a cup of oil. The horizontal arms are connected with a vertical spindle, upon which is an endless screw, working in a toothed wheel connected with a train of wheels, furnished with indices capable of registering one mile and decimal multiples of a mile up to 1,000 miles. A pinion C upon the axis of one of the wheels (which, in the figure, occupies a place too high) acts in a rack J, drawing it upwards by the ordinary motion of the revolving cups. The rack is pressed to the pinion by a spring, and, when it has been drawn up, it can be pressed by hand in opposition to the spring so as to release it from the pinion, and can then be pushed down, again to be raised by the action of the wheel-work. The rack is connected at the bottom with a sliding rod D, which passes down into the chanlber below, where it draws up the sliding pencil-carrier E. The pencil F, which it carries, traces its indications upon the sheet of paper wrapped round a barrel, whose axis is vertical, and which by spindle connexion with the clock H is made to revolve in 24 hours. The revolving cups and wheel-work are so adjusted that a motion of the pencil upwards of one inch represents a motion of the air through 100 miles. The curve traced upon

ROBINSON'S ANEMOMETER; RAIN GAUGES. *tv*

the barrel exhibits, therefore, the aggregate of the air's movements, and also the air's velocity, at every instant of the day.

In the year 1860, on July 3, 4, and 13, experiments were made in Greenwich Park, with the instrument then in use, to ascertain the correctness of the theory of Robinson's anemometer; the point to be verified being that the scale of the instrument, founded on the supposition that the horizontal motion of the air is about three times the space described by the centers of the cups, is correct.

A post about 5 feet high with a vertical spindle in the top was erected, and on this spindle turned a horizontal arm, carrying at the extremity of its longer portion Robinson's anemometer, and on its shorter portion a counterpoise. The distance from the vertical spindle of the post to the vertical axis of the anemometer was 17ft, Sin·'7. The reading of the dial was taken, and then the arm was made to revolve in the horizontal plane 50 or 100 times, an attendant counting the number of revolutions, and the reading of the dial was again taken. In this manner 1,000 revolutions were made in the direction N.E.S.W.N., and 1,000 revolutions in the direction N.W.S.E.N. In some of' the experiments the air was sensibly quiet, and in others there was a little wind; the result was,

For a movement of the instrument through one mile,

The results from rapid revolutions and from slow revolutions were sensibly the same.

This may be considered as confirming in a very high degree the accuracy of the theory.

§ 23. *Rain Gauges.*

'The rain-gauge connected with Osler's anemometer is 50 feet S inches above the ground, and 205 feet 6 inches above the mean level of the sea. It exposes to the rain an area of 200 square inches (its horizontal dimensions being 10 by 20 inches).

The collected water passes through a tube into a vessel suspended in a frame by spiral springs, which lengthen as the water increases, until 0'24 of an inch is collected in the receiver; it then discharges itself by means of the following modification of the syphon. A copper tube, open at both ends, is fixed in the receiver, in a vertical position, with its end projecting below the bottom. Over the top of this tube a larger tube, closed at the top, is placed loosely. The smaller tube thus forms the longer leg, and the larger tube the shorter leg, of a syphon. The water, having risen to the top of the smaller tube, gradually falls through it into the uppermost portion of a tumbling bucket, fixed in a globe under the receiver. When ful1, the bucket falls over,

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throwing the water into a small pipe at the lower part of the globe; the water completely fills the bore of the pipe; its descent causes an imperfect vacuum in the globe. sufficient to cause a draught in the longer leg of the syphon, and the whole contents run off. After leaving the globe, the water is carried away by a waste-pipe attached to the building. The springs then shorten and raise the receiver. The ascent and descent of the water-vessel move a radius-bar which carries a pencil ; and this pencil makes a trace upon the paper carried by the sliding board of the selfregistering anemometer. As the trace is rather long in proportion to the length of the radius-bar, the bar has now been furnished by Mr. Browning with a "parallel motion," which makes the trace sensibly straight.

The scale of the printed paper was adjusted by repeatedly filling the water-vessel until it emptied itself, then weighing the water, and thus ascertaining its bulk, and dividing this bulk by the area of the surface of the rain receiver.

A second gauge, with an area 77 square inches nearly, is placed close to the preceding, the receiving surface of both being on the same horizontal plane.

A third gauge is placed on the roof of the Octagon room, at 38 feet $4\frac{1}{9}$ inches above the ground, and 193 feet $2\frac{1}{2}$ inches above the mean level of the sea. It is a simple cylinder gauge, 8 inches in diameter and about $50\frac{1}{4}$ square inches in area. The height of the cylinder is $13\frac{1}{2}$ inches; at the depth of 1 inch from the top within the cylinder is fixed a funnel (an inverted cone) of 6 inches perpendicular height; with the point of this funnel is connected a tube, $\frac{1}{5}$ of an inch in diameter, and $1\frac{1}{2}$ inch in length; $\frac{3}{2}$ of an inch of this tube is slightly curved, and the remaining $\frac{3}{2}$ of an inch is bent upwards, terminating in an aperture of $\frac{1}{8}$ of an inch in diameter. By this arrangement, the last few drops of water remain in the bent part of the tube, and the water is some days evaporating. The upper part of the funnel or bore of the cone is connected with a brass ring, which has been turned in a lathe, and this is connected with a circular piece 6 inches in depth, which passes outside the cylinder, and rests in a water joint, attached to the inner cylinder, and extending all round.

A fourth gauge is placed on the top of the Library; it is a funnel, whose top has a diameter of 6 inches; its exposed area is $28\frac{1}{4}$ square inches nearly. The receiving surface of the gauge is 22 feet 4 inches above the ground, and 177 feet 2 inches above the mean level of the sea.

A fifth gauge is planted on the roof of the Photographic Thermometer shed, 10 feet above the ground, and 164 feet 10 inches above the mean level of the sea. Its construction is the same as that of the third gauge.

A sixth gauge is a self-registering rain-gauge on Crosley's construction, made by Watkins and Hill. The surface exposed to the rain is 100 square inches. The collected water falls into a vibrating bucket, whose receiving concavity is entirely above the center of motion, and which is divided into two equal parts by a partition whose plane passes through the axis of motion. The pipe from the rain-receiver ter-

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RAIN GAUGES: ELECTRICAL APPARATUS, lvi

minates immediately above the axis. Thus that part of the concavity which is highest is always in the position for receiving water from the pipe. When a certain quantity of water has fallen into it, it preponderates, and, falling, discharges its water into a cistern below; then the other part of the concavity receives the rain, and after a time preponderates. Thus the bucket is kept in a state of vibration. To its axis is attached an anchor with pallets, which acts upon a toothed wheel by a process exactly the reverse of that of a clock-escapement. This wheel communicates motion to a train of wheels, each of which carries a hand upon a dial-plate; and thus inches, tenths, and hundredths are registered. Sometimes, when the escapement has obviously failed, the water which has descended to the lower cistern has again been passed through the gauge, in order to enable an assistant to observe the indication of the dial-plates without fear of an imperfection in the machinery escaping notice. The gauge is placed on the ground, 21 feet South of the Magnetic Observatory, and 156 feet 6 inches above the mean level of the sea.

The seventh and eighth gauges are placed near together, about 16 feet south of the Magnetic Observatory, 5 inches above the ground, and 155 feet 3 inches above the mean level of the sea. They are similar in construction and area to No.3. These cylinders are sunk about 8 inches in the ground.

All these gauges, except No. 7, are read at 22^h daily; in addition, Crosley's gauge and No. 8 are read daily at 9^h p.m., and No. 7 at the end of each month only, to check the summation of the daily readings of No. 8. All are read at midnight of the last day of each month.

Gauges Nos. 1, 2, 3, 5, 8 were Inade by Messrs. Negretti and Zambra; No.4 by Troughton ; No.6 by Watkins and Hill; and No. 7 is an old gauge.

§ 24. *Electrical Apparatus.*

The electrical apparatus consists of two parts, namely, the Moveable Apparatus, which is connected with a pole nearly 80 feet high planted 7 feet North and 2 feet East of the north-east angle of the north arm of the Magnetic Observatory (as extended in 1862); and the Fixed Apparatus, which is mounted in a projecting window in the ante-room of the Magnetic Observatory.

On the top of the pole is fixed a projecting cap, to which are fastened the ends of two iron rods, which terminate in a pit sunk in the ground, and are kept in tension by attached weights. These rods are to guide the moveable apparatus in its ascents and descents. Near the bottom of the pole is fixed a windlass; the rope upon which it acts passes over a pulley in the cap, and is used to raise the moveable apparatus, which when raised to the top is suspended on a hook.

GREENWICH MAGNETICAL AND METEOROLOGICAL OBSERVATIONS, 1870. *h*

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The moveable apparatus consists of the following parts *:-A* plank in a nearly vertical position is attached to perforated iron bars, which slide upon the iron rods. On the upper part of this plank is a cubical *box.* The box incloses a stout pillar of glass, having a conical hollow in· its lower part. In the bottom of the box there is a large hole through which a cone of copper passes into the conical hollow of the glass pillar. In the lower part of the box a gas-lamp is placed, by the flame of which the copper cone and the lower part of the glass pillar are kept in a state of warmth. A copper wire is fastened round the glass pillar; its end is carried to a similar glass pillar, warmed in the same manner, near the north-western turret of the Octagon room; by this wire, whose length is about 400 feet, the atmospheric electricity is collected. To this wire, near the box, is attached another copper wire (now covered with gutta percha) 0·1 inch in diameter, and about 73 feet long, at the end of which is a hook; a loaded brass lever connected with the fixed apparatus presses upon this hook, and thus keeps the wire in a state of tension, and at the same time establishes the electrical communication between the long horizontal wire and the fixed apparatus.

The fixed apparatus consists of these parts *:*-A glass bar, nearly 3 feet long, and thickest at its middle, is supported in a horizontal position, its ends being fixed in pieces of wood projecting downwards from the roof of the projecting window. Near to each end is placed a small gas-lamp, whose chimney encircles the glass, and whose heat keeps the glass in a state of warmth proper for insulation. A brass collar surrounds the center of the glass bar; it carries one brass rod, projecting vertically upwards through a hole in the roof of the window-recess, to which rod are attached a small metallic umbrella and the loaded lever above-mentioned; and it carries another rod projecting vertically downwards, to which is attached a horizontal brass tube in an East and West direction. On the North and South sides of this tube there project four horizontal rods, through the ends of which there pass vertical rods, which can be fixed by screws at any elevation; these are placed in connexion with the electrometers, which rest on the window seat.

The electrometers during the year 1870 consisted of two Volta's Electrometers, denoted by Nos. 1 and 2; a Henley's Electrometer; a Ronalds' Spark Measurer; a Dry-pile Apparatus; and a Galvanometer.

Volta 1 and Volta 2 are of the same construction; each is furnished with a pair of straws 2 Paris inches in length; those of the latter being much heavier than those of the former: each instrument is furnished with a graduated ivory scale, whose radius is 2 Paris inches, and it is graduated into half Paris lines. In the original construction of these instruments it was intended that each division of $No. 2$ should correspond to five of No.1: the actual relation between them has not yet been determined by observations at the Royal Observatory. The straws are suspended by hooks of fine

ELECTROMETERS.

copper wire to the suspension-piece, and they are separated by an interval of half a line.

Henley's Electrometer is supported on the West end of the large horizontal tube by means of a vertical rod fixed in it. On each side of the upper part of this rod is affixed a semicircular plate of ivory, whose circumference is graduated; at the centers of these ivory plates two pieces of brass are fixed, which are drilled to receive fine steel pivots, carrying a brass axis, into which the index or pendulum is inserted; the pendulum terminates with a pith ball. The relation between the graduations of this instrument and those of the other electrometers has not been determined. This instrument has seldom been affected till Volta 2 has risen to above 100 divisions of its scale.

The spark measurer consists of a vertical sliding rod terminated by a brass ball, which ball can be brought into contact with one of the vertical rods before referred to, also terminating in a ball; and it can be moved from it or towards it by means of a lever, with a wooden handle. During the operation of separating the balls, an index runs along a graduated scale, and exhibits the distance between the balls, and this distance measures the length of the spark.

The electrometers and the spark measurer were originally constructed under the superintendence of Francis Ronalds, Esq., but have since received small alterations.

The dry-pile apparatus was made by Watkins and Hill; it is placed in connexion with the brass bar by a system of wires and brass rods. The indicator, which vibrates between the two poles, is a small piece of gold leaf. 1'his instrument is very delicate, and it indicates at once the quality of the electricity. \Vhen the inclination of the gold leaf is such that it is directed towards the top of either pile, it remains there as long as the quantity of electricity continues the same or becomes greater: the position is sometimes expressed in the notes by the words" as far as possible." The angle which the gold leaf makes with the vertical at this time is about 40[°].

The galvanometer was made by Gourjon of Paris, and consists of an astatic needle, composed of two large sewing needles, suspended by a split silk fibre, one of the needles of the pair vibrating within a ring formed by $2,400$ coils of fine copper wire. The connexions of the two portions of wire forming these 2,400 coils are so arranged that it is possible to use a single system of $1,200$ coils of single wire, or a system of 1,200 coils of double wire, or a system of 2,400 coils of single wire: in practice the last has always been used. A small ball communicating by a wire with one end of the coils is placed in contact at pleasure with the electric conductor, and a wire leading from the other end of the coil communicates with the earth. An adjustible circular card, graduated to degrees, is placed .immediately below the upper needle; the numeration of its divisions proceeds in both directions from a zero. One of these directions is distinguished by the letter A , and the other by the letter B ; and the nature of the indication represented by the deflection of the needle towards A or towards B will be ascertained from the following experiment. A voltaic battery being-formed by means
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of a silver coin and a copper coin, having a piece of blotting paper moistened with saliva between them: when the copper touches the small ball, and the wire which usually communicates with the earth is made to touch· the silver, the needle turns towards A; when the silver touches the small ball, and the wire is made to touch the copper, the needle turns towards B.

§ 25. *Explanation cifthe Tables* of *Meteorological Observations.*

The mean daily value of the difference between dew-point temperature and airtemperature is the difference between the two numbers in the sixth and seventh columns. The Greatest and Least are the greatest and least among the differences corresponding to the times of observation in the civil day, or they are found from the absolute maxima and minima, as determined by comparing the observations of the self-registering wet-bulb thermometers with those of the self-registering dry-bulb thermometers.

The difference between the mean temperature for the day and the mean for the same day of the year on an average of fifty years, is found by comparison with a table of results deduced by Mr. Glaisher from fifty years' observations, made at the Royal Observatory, ending 1863.

Little explanation of the results deduced from Osler's Anemometer appears to be necessary. It may be understood generally that the greatest pressure occurred in gusts of short duration.

To 1867, October 31, the indication of Robinson's Anemometer was read off every day at 22^h (10^h A.M.), and the difference between consecutive readings was entered opposite to the civil day on which the first reading was taken. Fron1 1867, November 1, the daily values have been extracted from the sheets of the continuous record, applying to the interval from midnight to midnight, and are entered opposite to the civil day to which each value belongs.

The daily register of rain is given for each civil day ending at midnight. This applies to the Cylinder Rain-gauge partly sunk in the ground, described above as the " eighth."

For understanding the divisions of time under the heads of Electricity and Weather, the following remarks are necessary:—The day is divided by columns into two parts (from midnight to noon, and from noon to nlidnight), and each of these parts is roughly subdivided into two or three parts by colons (\cdot) . Thus, when there is a single colon in the first column, it denotes that the remarks before it apply (roughly) to the interval from midnight to 6 A.M., and those following it to the interval from 6 A.M. to noon. When there are two colons in the first column, it is to be understood that the twelve hours are divided into three nearly equal parts of four hours each. And similarly for the second column.

TABLES OF METEOROLOGICAL OBSERVATIONS:
METEOROLOGICAL NOTATION. lxi $\frac{1}{2}$ and $\frac{1}{2}$ $\frac{1}{2}$, $\frac{1$

The following is the explanation of the notation employed for record of electrical observations, it being premised that the quality of the Electricity is always to be supposed positive when no indication of quality is given:-

The duplication of the letter denotes an intensity of the modification described, thus, s s is very strong; v v, very variable.

The Clouds and Weather are described generally by Howard's Nomenclature; the figure denotes the proportion of sky covered by clouds, the whole sky being represented by 10. The notation is as follows:

$l**xii**$ INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

The foot-notes show the means and extremes of readings, and their departure in each month from average values, as found from the preceding Twenty-nine Years' Observations; those relating to Humidity have been calculated from the Fifth Edition of Glaisher's Hygrometrical Tables.

§ 26. *Observations* cif *Luminous Meteors.*

In arranging for the observations of meteors, the directions circulated by the Committee of the British Association have received the most careful attention. The observers have been educated in the knowledge of the principal stars by observations of the stars themselves, and by means of globes and maps. The general instruction to all observers has been, to look out for meteors on every clear night; but the observer specially appointed for the evening's duties has been more particularly charged with this observation.

On the nights specially mentioned in the directions of the British Association Committee, greater attention was given to the sky, and the observations of meteors were made more systematically. The principal nights are, January 2 and 10; February 6; March 1; April 19; May 18; June 6 and 20; July 17, 20, and 29; August 3, August 7-13; September 10; October 1 and 23; November 9-14, November 19, 28, and 30; December 8-14, especially December 11. A more extended list of days has been published by the British Association Committee.

Special arrangements were made in the August period for observing till the morning; and in the November period for observing through the night, one or two observers being on duty till midnight, and then all the observers till daybreak. The observers were so stationed as to command different views of the sky, to secure observation of all the meteors which might present themselves, and to guard against the ohservation of the same meteor by different observers.

The observers in the year 1870 were Mr. Nash, Mr. Wright, Mr. Schultz, Mr. Marriott, Mr. W. Bishop, and Mr. W. Barber. Their observations are distinguished by the initials N., W., S., M., W. B., and E., respectively.

§ 27. *Details of the Chemical Operations for the Photographic Records.*

Mr. Glaisher has drawn up the following account of the Chemical Processes employed in the Photographic Operations for the self-registration of the Magnetical and Meteorological Indications.

LUMINOUS METEORS: PRIMARY PHOTOGRAPHY. *lxiii*

CHEMICAL PREPARATION AND TREATMENT OF THE PHOTOGRAPHIC PAPER FOR PRIMARIES.

The paper used in 1870 is principally furnished by Hollingsworth and Towgood; it is strong and of even texture, and is prepared expressly for Photographic purposes.

First Operation.-Preliminary Preparation of the Paper.

The chemical solutions used in this process are the following :-

(1.) Sixteen grains of Iodide of Potassium are dissolved in one ounce of distilled water.

(2.) Twenty-four grains of Bromide of Potassium are dissolved in one ounce of distilled water.

(3.) When the crystals are dissolved, the two solutions are mixed together, forming the iodising solution. The mixture will keep through any length of time. Immediately before use, it is filtered through filtering paper.

A quantity of the paper, sufficient for the consumption of several weeks, is treated in the following manner, sheet after sheet.

The sheet of paper is pinned by its four corners to a horizontal board. Upon the paper, a sufficient quantity (about 50 minims, or $\frac{5}{48}$ of an ounce troy) of the iodising solution is. applied, by pouring it upon the paper in front of a glass rod, which is then moved to and fro till the whole surface is uniformly wetted by the solution. Or, the solution may be evenly distributed by means of a camel-hair brush.

The paper thus prepared is allowed to remain in a horizontal position for a few minutes, and *is* then hung up to dry in the air; when dry, it is placed in a drawer, and may be kept through any length of time.

Second Operalion.-Rendering the Paper sensitive to the Action of *Light.*

A solution of Nitrate of Silver is prepared by dissolving 50 grains of crystallized Nitrate of Silver in one ounce of distilled water. Since the magnetic basement has been used for photography, 15 grains of Acetic Acid have always been added to the solution.

Then the following operation is performed in a room illuminated by yellow light.

The paper is pinned as before upon a board somewhat smaller than itself, and (by means of a glass rod, as before,) its surface is wetted with 50 minims of the Nitrate of Silver solution. It is allowed to remain a short time in a horizontal position, and, if any part of the paper still shines from the presence of a part of the solution unabsorbed into its texture, the superfluous fluid is taken off by the application of blotting paper.

lxiv INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

The paper, still damp, is immediately placed upon the cylinder, and is covered by the exterior glass tube, and the cylinder is mounted upon the revolving apparatus, to receive the spot of light formed by the mirror, which is carried by the magnet; or to receive the line of light passing through the thermometer tube.

Third Operation.-Developmen:t qf *the Photographic Trace.*

When the paper is removed from the cylinder, it is placed as before upon a board, and a saturated solution of Gallic Acid, to which a few drops of Aceto-Nitrate of Silver are occasionally added, is spread over the paper by means of a glass rod, and this action is continued until, the trace is fully developed. The solutions are kept in the magnetic basement, and are always used at the temperature of that room. When the trace is well developed, the paper is placed in a vessel with water, and repeatedly washed with several waters; a brush being passed lightly over both sides of the paper to remove any crystalline deposit.

Fourth Operation.-Fixing the Photographic Trace.

The Photograph is placed in a solution of Hyposulphite of Soda, made by dissolving four or five ounces of the Hyposulphite in a pint of water; it is plunged completely in the liquid, and allowed to remain from one to two hours, until the yellow tint of the Iodide of Silver is removed, After this the sheet is washed repeatedly with water, allowed to remain immersed in water for 24 hours, and afterwards placed within folds of cotton cloths till nearly dry. Finally it is placed between sheets' of blotting-paper, and is pressed.

CHEMICAL PREPARATION AND TREATMENT OF THE PHOTOGRAPHIC PAPER FOR SECONDARIES.

Before taking a Secondary, the Primary is examined to ascertain whether the tint of the photographic curve is sufficiently dark. If it is not, the Primary is laid, face downwards, upon a desk of transparent plate-glass, below which is a large silvered plane mirror, so placed that the light from the sky is reflected upwards through the transparent glass and through the Primary;' and the photographic curve is seen from the upper side or back with perfect distinctness. An assistant then darkens the back of the photographic curve by the application of sepia; the original photograph being untouched.

The paper used for the Secondaries is made by Rive; it is a strong wove paper, of tolerably even texture, thin, but able to bear a great deal of wear.

PRIMARY AND SECONDARY PHOTOGRAPHY. lxv

First Operation.-Preliminary Preparation of the Paper.

The chemical solution required for this purpose is as follows :—

Two grains of Chloride of Ammonium are dissolved in one ounce of distilled water. A sufficient quantity of this solution is placed in a flat-bottomed porcelain dish, and sheets of paper, one by one, are plunged within it; care being taken that no air bubbles remain between the paper and the solution; this may be prevented by slight pressure over the sheet by means of a bent glass rod. When a few sheets are thus immersed, they are turned over, and are taken out and hung to dry. Any number of sheets may thus be prepared.

An equally good result is obtained, by spreading over one side by means of a glass rod, as in the preparation of the Primaries, a solution of Chloride of Ammonium made by dissolving five grains of the chloride in one ounce of distilled water.

Second Operation.—Rendering the Paper sensitive to the Action of Light.

The solution required for this purpose is as follows :-

To a filtered solution of Nitrate of Silver (made by dissolving 50 grains of Crystallized Nitrate of Silver in one ounce of distilled water) some strong solution of Ammonia is added; the whole becomes at first of a dark brown colour, but when a sufficient quantity of Ammonia is added the solution becomes perfectly clear; a few crystals of Nitrate of Silver are then added till the solution is a little dull, forming "Ammoniacal Nitrate of Silver"; it is then ready for use.

The following operation is performed in a room illuminated by yellow light :-

By means of a glass rod this solution is spread over the paper, whilst pinned on a board; the paper is dried before a fire, and is then in a fit state to be used for producing a Secondary.

Third Operation.-Formation of *the Photographic Copy.*

A sheet of the paper so prepared is placed in a printing fralne with its prepared side upwards, upon a bed of blotting paper resting upon a sheet of plate-glass; the Primary is then placed on the paper with its own face downwards; and as it is necessary, for obtaining a correct copy of the Primary, that it should be in close contact with the prepared surface, a second sheet of plate-glass is placed over it, and the two are pressed together by clamps and screws. The whole is then exposed to the light (the Primary to be copied being above the paper on which the copy is to be made). The time required to produce a copy depends, in a great measure, upon the thickness of the paper on which the Primary is made, and on the actinic quality of the light; a period of five minutes in a bright sunshine, or one hour in clear daylight, is generally sufficient.

GREENWICH MAGNETICAL AND METEOROLOGICAL OBSERVATIONS, 1870.

l.rvi INTRODUCTION TO GREENWICH MAGNETICAL OBSERVATIONS, 1870.

Fourth Operation.—Fixing the Photographic Secondary.

When an impression has been thus obtained, it is necessary that the undecomposed Salts of Silver remaining in the paper he removed.

:For this purpose the Secondary is at once plunged into water and well washed on both sides, passing a camel-hair brush over every part of it; it is then plunged into a solution of Hyposulphite of Soda (made by dissolving two or three ounces of the Hyposulphite in a pint of water), and is left through a period varying from half an hour to an hour. It is then removed, and washed in plain water several times; and running water is allowed to pass over it for twenty-four hours.

The sheets are then placed within the folds of drying cloths, till nearly dry, and finally between sheets of blotting paper.

The process of obtaining a Tertiary from a Secondary is in every respect the same as that of obtaining a Secondary from a Primary.

§ 28. *Personal Establisllment.*

The personal establishment during the year 1870 has consisted of James Glaisher, Esq., F.R.S., Superintendent of the Magnetical and Meteorological Department, and Mr. William Carpenter Nash, Assistant.

Three or four computers have usually been attached to the Department.

Royal Observatory, Greenwich, 1871, November 10.

G. B. AIRY.

ROYAL OBSERVATORY, GREENWICH.

RESULTS

OF

MAGNETICAL 0 BSERV **A.TIONS.**

1870.

GREENWICH OBSERVATIONS, 1870.

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 $\label{eq:1} \widetilde{Q}^{(n)} = \frac{1}{n} \sum_{i=1}^n \frac{1}{n_i} \sum_{j=1}^n \frac{1}{n_j} \sum_{j=1}$ \sim $\mathcal{L}(\mathcal{A})$. $\sim 10^{-10}$

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ROYAL OBSERVATORY, GREENWICH.

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TABLE V.-MEAN MONTHLY DETERMINATION of the HORIZONTAL MAGNETIC FORCE (diminished by a Constant 0.8600 nearly), uncorrected for TEMPERATURE, at every HOUR of the DAY; obtained by taking the MEAN of all the DETERMINATIONS at the same HOUR of the DAY through each MONTH.

The Thermometer on the box inclosing the Horizontal Force Magnetometer was read generally nine times every day. The means of the readings taken for the same nominal hour through each month show no sensible Mean Diurnal Ine

TABLE VI.

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TABLE VII.-MEAN VERTICAL MAGNETIC FORCE (diminished by a Constant 0'9600 nearly) on each ASTRONOMICAL DAY, as deduced from

TABLE VIII.-MEAN MONTHLY DETERMINATION of the VERTICAL MAGNETIC FORCE (diminished by a Constant 0'9600 nearly) uncorrected for TEMPERATURE, at every HOUR of the DAY; obtained by taking the MEAN of all the DETERMINATIONS at the same HOUR of the DAY through each MONTH.

The Thermometer on the box inclosing the Vertical Force Magnetometer was read generally nine times every day. The means of the same nominal hour through each month show no sensible Mean Diurnal Inequality of Temperature.

ROYAL OBSERVATORY, GREENWICH.

INDICATIONS

MAGNETOMETERS

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ON SIXTEEN DAYS OF GREAT MAGNETIC DISTURBANCE.

1870.

GREENWICH OBSERVATIONS, 1870. **B**

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INDICATIONS OF THE MAGNETOMETERS

The Symbol : attached to a time denotes that the reading will apply equally well to a considerable range of time near that which is recorded. A brace denotes that at this time the curve of the Vertical Force was dislocated, and the difference of the numbers included by the brace shows the amount of the displacement,

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

February 1. The spot of light for Horizontal Force was off the sheet in the direction of *decreasing* force from 10^h. 27^m. to 11^h. 15^m.

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INDICATIONS OF THE MAGNETOMETERS

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by the brace shows the amount of the displacement.

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

The indications are taken from the sheets of the Photographic Record, except where an asterisk is attached to the number, in which instances they are inferred from
observations made with the telescope in the ancient manner

the numbers included by the brace shows the amount of the displacement.
April 5. The spot of light for Horizontal Force was off the sheet in the direction of *decreasing* force from 2^h. 49^m, to 3^h, 0^m.

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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(xvi) INDICATIONS OF THE MAGNETOMETERS

The indications are taken from the sheets of the Photographic Record, except where an asterisk is attached to the number, in which instances they are inferred from
observations made with the telescope in the ancient manne

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

GREENWICH OBSERVATIONS, 1870.

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

by the brace shows the amount of the displacement.

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

September 24. The spot of light for Horizontal Force was off the sheet in the direction of *decreasing* force from 12^h . 2^m , to 14^h , 42^m .;
from 14^h , 53^m , to 16^h , 28^m ,; and from 16^h , 53^m , to 2

GREENWICH OBSERVATIONS, 1870.

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INDICATIONS OF THE MAGNETOMETERS

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

GREENWICH OBSERVATIONS, 1870.

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INDICATIONS OF THE MAGNETOMETERS

October 25. The spot of light or Declination was off the sheet in the direction of *increasing* declination from 3^{h} , $3^{6\text{m}}$, to 3^{h} , 59^{m}

The indications are taken from the sheets of the Photographic Record, except where an asterisk is attached to the number, in which instances they are interest from the sheets of the Photographic Record, except where an as of the numbers included by the brace shows the amount of the displacement.

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

November 4. VERTICAL FORCE.—The adjustments were altered, so that the readings were increased by 19^{div}34 or by 0'01074 parts of the whole Vertical Force. It will be necessary therefore to diminish the indications on Nove

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INDICATIONS OF THE MAGNETOMETERS

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For the Horizontal and Vertical Forces, increasing readings denote increasing forces.

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INDICATIONS OF THE MAGNETOMETERS

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 $\label{eq:1} \mathbf{X} = \frac{1}{2} \sum_{i=1}^{N} \frac{1}{2} \sum_{j=1}^{N} \frac$

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 $\label{eq:2.1} \frac{1}{\sqrt{2\pi}}\int_{\mathbb{R}^3}\frac{d\mu}{\sqrt{2\pi}}\left(\frac{d\mu}{\mu}\right)^2\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\$

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ROYAL OBSERVATORY, GREENWICH.

RESULTS

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MAGNETIC DIP.

1870.

GREENWICH OBSERVATIONS, 1870.

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The initial N is that of Mr. W. C. Nash.

In the month of January several observations were made with an instrument by Dover, of the Kew pattern, provided with needles 3 inches in length, marked A I and A 2; and observations were also made in the months January to

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MONTHLY AND YEARLY MEANS OF MAGNETIC DIPS,

For this table the monthly means have been formed without reference to the hour at which the observation was made on each day. In combining the monthly results, to form the annual means, weights have been given proportional to the number of observations.

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RESULTS of OBSERVATIONS of MAGNETIC DIP at the Hours of Observation 9^h. a.m. and 3^h. p.m.

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 $\mathcal{L}(\mathcal{L})$ and $\mathcal{L}(\mathcal{L})$ $\label{eq:2.1} \frac{1}{\sqrt{2\pi}}\int_{\mathbb{R}^3}\frac{d\mu}{\sqrt{2\pi}}\left(\frac{d\mu}{\mu}\right)^2\frac{d\mu}{\sqrt{2\pi}}\int_{\mathbb{R}^3}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\$ $\label{eq:2} \frac{1}{2} \sum_{i=1}^n \frac{1}{2} \sum_{j=1}^n \frac{1}{$ $\label{eq:2.1} \frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac{1}{\sqrt{2}}\sum_{i=1}^n\frac$ $\mathcal{L}^{\text{max}}_{\text{max}}$, $\mathcal{L}^{\text{max}}_{\text{max}}$ \sim \sim $\label{eq:2.1} \frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\right)^{2\pi} \frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\$

 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\left(\frac{1}{\sqrt{2}}\right)^2\left(\frac{1}{\sqrt{2}}\right)^2\left(\frac{1}{\sqrt{2}}\right)^2\left(\frac{1}{\sqrt{2}}\right)^2\left(\frac{1}{\sqrt{2}}\right)^2.$

ROYAL OBSERVATORY, GREENWICH.

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OBSERVATIONS

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DEFLEXION OF A MAGNET

FOR

ABSOLUTE MEASURE

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1870.

(xlviii) OBSERVATIONS AND COMPUTATIONS OF DEFLEXION OF A MAGNET FOR ABSOLUTE MEASURE OF HORIZONTAL FORCE,

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The position of the Deflecting Magnet with regard to the suspended Magnet is always that which was formerly termed "Lateral." The Deflecting Magnet is placed on the East side of the suspended Magnet, with its marked pole a The lengths of ι foot and $\iota \cdot \iota$ foot answer to 304.8 and 396.2 millimètres respectively.

The initial N is that of Mr. W. C. Nash.

In the following calculations every observation is reduced to the temperature 35° .

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 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2.$

 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2.$

 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2.$ $\label{eq:2.1} \frac{1}{\sqrt{2\pi}}\int_{\mathbb{R}^3}\frac{d\mu}{\sqrt{2\pi}}\left(\frac{d\mu}{\mu}\right)^2\frac{d\mu}{\sqrt{2\pi}}\int_{\mathbb{R}^3}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\sqrt{2\pi}}\frac{d\mu}{\$

 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2.$

 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\right)^{2\alpha} \frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\right)^{\alpha} \frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}$

ROYAL OBSERVATORY, GREENWICH.

RESULTS

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METEOROLOGICAL OBSERVATIONS.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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The highest in the month was 47^o 1 on the 7th ; and the lowest was 20° 9 on the 20th.
The mean
Elastic Force of Vapour. The mean for the month was $0^{4\circ}$, being $0^{5\circ}$ of lower than the average of the preceding

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 6.5 .

Ozone.
The mean amount for the month, on a scale ranging from 0 to 10, was 2'2.

Wind.

The proportions were of N. 5, S. 10, W. 8, E. 6, and Calm 2. The greatest pressure in the month was more than 3^{obs} on the square foot on the 8th. RAIN.

Fell on 15 days in the month, amounting to i^{in} 49, as measured in the simple cylinder gauge partly sunk below the ground; being o^{in} 38 *less* than the average fall of the preceding 55 years.

RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

BAROMETER READINGS FROM EYE-OBSERVATIONS.

The first maximum in the month was $29^{\ln} \cdot 863$ on the 1st; the first minimum in the month was $29^{\ln} \cdot 545$ on the 2nd.

The second maximum in the month was $29^{\ln} \cdot 648$ on the 3rd; the second minimum in the month w The absolute maximum ,, was $30^{10} \cdot 202$ on the 12th; the fourth minimum ,, was $30^{10} \cdot 108$ on the 15th; the fifth maximum , was $30^{10} \cdot 971$ on the 20th; the fifth minimum ,, was 7 he sixth maximum , was 29^{10} was 29^{ln} , 700 on the 13th.
was 29^{ln} , 730 on the 13th.
was 29^{ln} , 395 on the 21st.
was 29^{ln} , 303 on the 24th.
was 29^{ln} , 303 on the 26th.

TEMPERATURE OF THE AIR.

The highest in the month was 55° 6 on the 28th; the lowest was 19° 4 on the 11th.
The range and the month was 36° 2.
The mean call the highest daily readings was 41° 4, being 4° 1 lower than the a

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HUMIDITY OF THE AIR.

Temperature of the Dew Point.

The highest in the month was 46° , 3 on the 28th; and the lowest was 10°, 2 on the 12th.

The mean γ , was 29° , being 5° , 4 lower than the average of the preceding 29 years.
 Weight of Vapour,—The mean f

CLOUDS.

The mean amount for the month, a clear sky being represented by 0 and a cloudy sky by *10,* was 7' 4.

Ozone.
The mean amount for the month, on a scale ranging from 0 to 10, was 3'5.

WIND.

The proportions were of N. 5, S. 10, W. 5, E. 8, and Calm o. The greatest pressure in the month was more than 30^{lbs} on the square foot on the 13th and 14th. RAIN.

Fell on 13 days in the month, amounting to o^{ia} 54, as measured in the simple cylinder gauge partly sunk below the ground; being I^{in} o3 less than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was $7 \cdot 5$.

OZONE.

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The mean amount for the month, on a scale ranging from \circ to 10, was 2° 9.

WIND.
The proportions were of N. 13, S. 5, W. 6, E. 7, and Calm 0. The greatest pressure in the month was 30^{1bs}'0 on the square foot on the 5th.
RAIN.

Fell on 11 days in the month, amounting to 2^{\ln} ° 05, as measured in the simple cylinder gauge partly sunk below the ground; being 0^{\ln} 46 *greater* than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

BAROMETER READINGS FROM ETE-OBSERVATIONS.

The absolute maximum in the month was 30^{in} 355 on the 4th; the first minimum in the month was 30^{in} 033 on the rst. The second maximum (3), was $30^{10} \cdot 333$ on the 16th; the absolute minimum (3), was $20^{10} \cdot 314$ on the 9th. was 29^{in} 829 on the 20th. was 30^{1n} · 238 on the 24th; the third minimum The third maximum $, \, ,$

was 29in . 994 on the 27th; the fourth minimum

 $\,$, $\,$

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was 29ⁱⁿ 886 on the 26th.

The fourth maximum $\ddot{}$ The range in the month was 1^{in} . 041.

The mean for the month was 29th . 984, being oth 218 higher than the average of the preceding 29 years.

TEMPERATURE OF THE AIR.

The highest in the month was 78° 7 on the 20th; the lowest was 26° o on the 4th.
The range , was 52° 7.

 \rightarrow The range

of all the highest daily readings was 62° o, being 4° 3 higher than the average of the preceding 29 years.
of all the lowest daily readings was 38° 4, being \circ 8 lower than the average of the preceding 29 years The mean $, \, ,$

The \mathtt{mean} \rightarrow \rightarrow

The mean daily range was 23° 6, being 5° 2 greater than the average of the preceding 29 years. The mean for the month was 48°'9, being 1°'8 higher than the average of the preceding 29 years.

HUMIDITY OF THE AIR.

Temperature of the Dew Point.

The highest in the month was 51° o on the 21st; and the lowest was 29° ; on the 1st.
The mean
Elastic Force of Vapour.—The mean for the month was 39° and the average of the preceding 29 years.
Weight of Vapo

Weight of *a Cubic Foot of Air.-*The mean for the month was 546 grains, being 3 grains *greater* than the average of the preceding 29 years.

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 4'0.

OZONE. The mean amount for the month, on a scale ranging from 0 to *10,* was 3' I.

WIND.
The proportions were of N. 5, S. 7, W. 11, E. 5, and Calm 2. The greatest pressure in the month was 9^{1bs}'1 on the square foot on the 10th.

RAIN.
Fell on 6 days in the month, amounting to o^{in} 28, as measured in the simple cylinder gauge partly sunk below the ground; being I^{in} 45 *less* than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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HUMIDITY OF THE AIR.

Temperature of the Dew Point.

The highest in the month was 55° on the aoth; and the lowest was 32° on the 4th.
The mean \qquad , was 45° r, being \circ o' 5 lower than the average of the preceding 29 years.
Elastic Force of Vapour.—The m

Weight of Vapour in a Cubic Foot of Air.—The mean for the month was 3^{gr*} . being 0^{gr*} I less than the average of the preceding 29 years.
Degree of Humidity.—The mean for the month was 73 (that of Saturation being repr

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 5.6 .

OZONE. The mean amount for the month, on a scale ranging from \circ to 10, was 3.6 .

WIND.

The proportions were of N. 6, S. 8, W. 9, E. 7, and Calm I. The greatest pressure in the month was 30^{bs} on the square foot on the 12th and 13th.

RAIN. Fell on 5 days in the month, amounting to oⁱⁿ⁴7, as measured in the simple cylinder gauge partly sunk below the ground; being 1¹ⁿ¹70 less than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

BAROMETER READINGS FROM EYE-OBSERVATIONS.

The absolute maximum in the month was 30^{in} · 367 on the 6th; the absolute minimum in the month was 29^{in} · 550 on the 10th. was $30^{11} \cdot 648$ on the 12th; the second minimum cases were $\frac{1}{2}$, was $20^{11} \cdot 715$ on the 17th. The second maximum ,, was $30^{\text{in}} \cdot 191$ on the 21st; the third minimum was 29in 771 on the 24th. The third maximum \mathbf{r} $\bar{\boldsymbol{\theta}}$ was 29^{in} 764 on the 27th. The fourth maximum was $29^{in} \cdot 973$ on the 25^{th} ; the fourth minimum $\overline{\boldsymbol{\beta}}$ $\,$, $\,$ was 29^{in} 944 on the 28th. The fifth maximum $\overline{}$

The range in the month was o^{in} 817.

The mean for the month was $29^{\text{in}} \cdot 947$, being $0^{\text{in}} \cdot 136$ higher than the average of the preceding 29 years.

TEMPERATURE OF THE AIR.

The highest in the month was 90° a on the 22nd; the lowest was 41° 4 on the 6th.

was $48^\circ \cdot 8$. The range $\,$, $\,$

The mean ,, of all the highest daily readings was $7^{0.8}$, being $3^{0.7}$ higher than the average of the preceding 29 years.
The mean ,, of all the lowest daily readings was $50^{0.7}$, being $0^{0.7}$ higher than the aver

The mean for the month was 60° g, being r° g higher than the average of the preceding 29 years.

Temperature of the Dew Point.

The highest in the month was 63° o on the 22nd; and the lowest was 41° on the 25th.

The mean $\begin{array}{c}$, was $50^{\circ} \cdot 6$, being $0^{\circ} \cdot 1$ lower than the average of the preceding 29 years.
Elastic Force of Vapour.—The mean for the month was $0^{\text{in}} \cdot 369$, being $0^{\text{in}} \cdot 303$ less than the average of

Weight of Vapour in a Cubic Foot of Air.—The mean for the month was 45"' I, being 05"' I less than the average of the preceding 29 years.
Degree of Humidity.—The mean for the month was 68 (that of Saturation being represen

Weight of a Cubic Foot of Air.-The mean for the month was 532 grains, being *the same as* the average of the preceding 29 years.

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 6.4.

Ozone.
The mean amount for the month, on a scale ranging from 0 to 10, was 3°2.

WIND. The proportions were of N. 8, S. 5, W. 13, E. 4, and Calm o. The greatest pressure in the month was 18^{1bs} \cdot 5 on the square foot on the 11th.

RAIN.

Fell on 4 days in the month, amounting to o^{in} 39, as measured in the simple cylinder gauge partly sunk below the ground; being i^{in} 55 less than the average fall of the preceding 55 years.

(lxiii)

RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

BAROMETER READINGS FROM EYE-OBSERVATIONS.

The first minimum in the month was $29^{1n} \cdot 653$ on the 4th. was 29^{in} + 497 on the 1 rth.
was 29^{in} + 689 on the 16th. The first maximum in the month was $29^{in} \cdot 966$ on the $7th$; the absolute minimum,

The second maximum ,, The absolute maximum The fourth maximum

was $29^{1n} \cdot 917$ on the 14th; the third minimum

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was $30^{\text{in}} \cdot 134$ on the 20th; the fourth minimum was 30^{1n} or 7 on the 28th; the fifth minimum

 $\;$, $\;$ was $29^{in.643}$ on the 25th. $\, \cdots \,$

was 29^{in} 555 on the 31st. \mathbf{r}

The range in the month was o^{in} \cdot 637.

The mean for the month was 29th 818, being oth of a higher than the average of the preceding 29 years.

TEMPERATURE OF THE AIR.

The highest in the month was 89° 7 on the 8th; the lowest was 44° 8 on the 2nd.

was 44° . 9. The range $\overline{}$

The mean of all the highest daily readings was 78° I, being 4° I higher than the average of the preceding 29 years. $, \, ,$

of all the lowest daily readings was 56° o, being 3° o higher than the average of the preceding 29 years. The mean $, \, , \,$

The mean daily range was 22°'1, being 1°'0 greater than the average of the preceding 29 years.

The mean for the month was 65° 4, being 3° 5 higher than the average of the preceding 29 years.

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HUMIDITY OF THE AIR.

Temperature of the Dew Point.

Tenepature of the Dew Point.

The highest in the month was 64° o on the oth; and the lowest was 47° on the and.

The mean the month was 55° o, being x° ; if sighter than t

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 6.6 . OZONE.

The mean amount for the month, on a scale ranging from \circ to 10, was $3'$ 4.

WIND.
The proportions were of N. 5, S. 7, W. 12, E. 7, and Calm o. The greatest pressure in the month was $10^{\text{hs}} \cdot 4$ on the square foot on the 5th.

RAIN.
Fell on 10 days in the month, amounting to 2^{\ln} or, as measured in the simple cylinder gauge partly sunk below the ground; being o^{\ln} 55 less than the average fall of the preceding 55 years.
RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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HUMIDITY OF THE AIR. *Temperature of tlte Dew Point.*

The highest in the month was 66° I on the (st; and the lowest was 41° γ on the 29th.

The mean
Elastic Force of Vapour.—The mean for the month was 5^{20} o, being 10^{18} of the preceding 29 years,
Weight of Vapour in a Cubic Foot of Air.—The mean for the month was $e^{in} \cdot 388$, being $e^{in} \cdot 292$ less t

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 6.2.

Ozone.
The mean amount for the month, on a scale ranging from 0 to 10, was 3'2.

WIND.

The proportions were of N. 13, S. 4, W. 7, E. 7, and Calm 0. The greatest pressure in the month was 22^{1b3} on the square foot on the 28th.
RAIN.

Fell on 10 days in the month, amounting to 2^{iu} '02, as measured in the simple cylinder gauge partly sunk below the ground; being 0ⁱⁿ 37 *less* than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

BAROMETER READINGS FROM EYE-OBSERVATIONS.

The first minimum in the month was $29^{\text{ln}} \cdot 325$ on the 2nd,

; the absolute minimum , was $29^{\text{ln}} \cdot 160$ on the 7th,

; the third minimum , was $29^{\text{ln}} \cdot 239$ on the 9th,

; the forth minimum , was $39^{\text{ln}} \cdot 669$ The first maximum in the month was $29^{\ln} \cdot 846$ on the 4th; the absolute minimum ,
The second maximum , was $29^{\ln} \cdot 601$ on the 8th; the third minimum ,
The third maximum , was $30^{\ln} \cdot 015$ on the 11th; the fourth m

The range in the month was $I^{in.209}$. The mean for the month was z9ⁱⁿ . 906, being oth og5 higher than the average of the preceding 29 years.

TEMPERATURE OF THE AIR.

The highest in the month was 72° 6 on the 1st; the lowest was 37° 4 on the 25th.
The range , was 35° 2.

of all the highest daily readings was $66^\circ \cdot 8$, being 1° \circ lower than the average of the preceding 29 years.

The mean of all the lowest daily readings was 46° 4, being 2° 9 lower than the average of the preceding 29 years. $, \, ,$

The mean The mean daily range was 20° 4, being 1° 9 greater than the average of the preceding 29 years.
The mean for the month was 55° 7, being 1° 6 lower than the average of the preceding 29 years.

HUMIDITY OF THE AIR.

Temperature of the Dew Point.

Temperature of the Dew Font.
The highest in the month was 61° i on the 2nd; and the lowest was 42° ; 3 on the 15th.
The mean
Elastic Force of Vapour.—The mean for the month was o^{10} ; 3 or the preceding 29 year

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 4'3.

OZONE. The mean amount for the month, on a scale ranging from o to 10, was 3'7.

WIND.

The proportions were of N. 4, S. 7, W. 9, E. 9, and Calm 1. The greatest pressure in the month was 30^{lbs} on the square foot on the 10th.

RAIN.
Fell on 8 days in the month, amounting to i^{in} 63, as measured in the simple cylinder gauge partly sunk below the ground; being o^{in} 80 less than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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The mean daily range was 15° , the state of the preceding 29 years.
The mean daily range was 15° , being 0° ; *preater* than the average of the preceding 29 years.
The mean for the month was 49° .8, being

HUMIDITY OF THE AIR.

Temperature of the Dew Point.

Temperature of the Deta Fount.
The highest in the month was 55° 9 on the 2nd; and the lowest was 33° 4 on the 11th.
The mean or the month was 55° 9 on the 2nd; and the lowest was 33° 4 on the 11th.
El

CLOUDS. The mean amount for the month, a clear sky being represented by o and a cloudy sky by io , was o^2

WIND.
The proportions were of N. 5, S. 8, W. 13, E. 5, and Calm o. The greatest pressure in the month was more than 30^{ths} o on the square foot on the 12th and 13th. RAIN.

Fell on 16 days in the month, amounting to 3^{in} 34, as measured in the simple cylinder gauge partly sunk below the ground; being o^{in} 57 greater than the average fall of the preceding 55 years.

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RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

The mean ,, was 38° , being $1^\circ \cdot 2$ *lower* than the average of the preceding 29 years.

Elastic Force of Vapour.-The mean for the month was o^{in} 233, being o^{in} on 6 less than the average of the preceding 29 years.

Weight of Vapour in a Cubic Foot of Air.-The mean for the month was 2grs · 7, being ogr · I less than the average of the preceding 29 years.

Degree of Humidity.-The mean for the month was 90 (that of Saturation being represented by 100), being 2 greater than the average of the precediug 29 years.

Weight of a Cubic Foot of Air.-The mean for the month was 548 grains, being the same as the average of the preceding 29 years.

CLOUDS.

The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was 6'1.

WIND.

The proportions were N. 5, S. 9, W. 10, E. 5, and Calm 1. The greatest pressure in the month was more than 30^{bs} o on the square foot on the 22nd and 24th. RAIN.

Fell on 9 days in the month, amounting to 1ⁱⁿ 20, as measured in the simple cylinder gauge partly sunk below the ground; being 1^{ta} 15 less than the average fall of the preceding 55 years.

$\left($ lxxiv $\right)$

RESULTS OF DAILY METEOROLOGICAL OBSERVATIONS

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The mean , was 29° s, being 7° 4 lower than the average of the preceding 29 years.
Elastic Force of Vapour.—The mean for the month was \circ ¹¹ 166, being \circ ¹¹ \circ 58 less than the average of the preceding 2

Weight of Vapour in a Cubic Foot of Air. The mean for the month was 1st 9, being ost 7 less than the average of the preceding 29 years.

Degree of Humidity.-The mean for the month was 86 (that of Saturation being represented by 100), being 2 less than the average of the preceding 29 years.

Weight of a Cubic Foot of Air. The mean for the month was 559 grains, being 7 grains greater than the average of the preceding 29 years.

CLOUDS. The mean amount for the month, a clear sky being represented by o and a cloudy sky by 10, was $8 \cdot r$.

WIND.
The proportions were of N. 11, S. 5, W. 5, E. 9, and Calm 1. The greatest pressure in the month was 22^{bs} o on the square foot on the 21st.

Fell on 22 days in the month, amounting to $3^{\text{ln}} \cdot 13$, as measured in the simple cylinder gauge partly sunk below the ground; being $1^{\text{ln}} \cdot 17$ greater than the average all of the preceding 55 years.

MAXIMA AND MINIMA BAROMETER-READINGS,

The following table contains the highest and lowest readings of the Barometer, reduced to 32° Fahrenheit, extracted from the photographic records. The readings are accurate; but the times are liable to great uncertainty, as the barometer frequently remains at its highest or lowest point through several hours. The time given is the middle of the stationary period. Where the symbol: follows the time, it denotes that the quicksilver has been sensibly stationary through a period of more than one hour.

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(Ixxviii) . ABSOLUTE MAXIMA AND MINIMA BAROMETER READINGS, AND MONTHLY METEOROLOGICAL MEANS,

	Mean Reading		TEMPERATURE OF THE AIR.											Mean		Mean		Mean Weight of	Mean additional
1870, MONTH.	of the Barometer.		Highest.		Range in Lowest. the Month.		Mean of all Mean of all Mean Daily the Highest.		the Lowest.	Range.		Mean Tempera- ture.		Tempera- ture of Dew Point.		Elastic Force of Vapour.		Vapour in a Cubic Foot of Air.	Weight required to saturate a Cubic Foot of Air.
January	in. 29.823		\bullet 50.9		\circ 19.6	\bullet 31.3	\bullet 42.6		\bullet 34°	\bullet 8.6		\bullet 38.3		\bullet 34'1	in.	0.196		grs. 2.3	grs. o. 5
February	29.693			55.6 19'4		36.2	41'4	31.9		a'5		36.2		29.7		0.165		1.0	o.6
March	29.865		61.1		23.1	38.0	46.9		34.0	12.9		39.6		34.7		0.301		2.3	\circ 5
April	29.984		78.7		26.o	52.7	62.0	38.4		23.6	48.9			$3q \cdot 2$	0.230		2.8		1'2
$May \ldots$	29.896		85.4		29.8	55.6	66.9	42° O		53.4 24.9			45.1	0.301		3.4		1'2	
June	29'947		90.5		41.4	48.8	74.8		50.7	24.1		60'g		50 [.] 6	o•36a		4.1		1.0
$July \dots \dots$	29.818		89.7		44.8	44.9	$78 \cdot I$		56.0 22'1			65.4		55.0		o*433		4.8	2.1
$August \dots$	29.804		81.0		41'	40.0	72.6		$53 \cdot 1$		19.5 61.1			52 ° 0	o*388		4.3		1.7
September.		29.906		72.6	37.4	35.2	66.8		46.4	20.4		55.7		50 · 5		o•367		4'1	0.8
October		29.570		68.6	32.4	36.2	58.2		42.8	15.4		49.8		45.3		0.303		3.5	o.6
November.		29.637		58.9	24.3	34.6	47.8		35.4	12.4		41.5		38.5		0.233		2.7	\circ 3
December	29.733		57.4		9.8	47.6	38.o		28.9	9'1		33.6		29.8		0.166	1.0		$\mathbf{0} \cdot \mathbf{3}$
$Means$	29.806			70.8 29'1		41.8	58.0	41'1		16.a		48.7		42.0 0.380		3.2		1'	
		Mean Degree			Mean	RAIN.				WIND.									
				Mean Weight			Amount			From From Osler's Anemometer. Robin-									
			of	of a		Amount Number	collected on											son's	
1870,		Humidity.		Cubic	of Cloud.	of	the Ground.			Calm or In Hours. Number of Hours of Prevalence of each Wind, Mean Daily referred to									Anemo- meter.
MONTH.		(Sat. $= 100.$		Foot of Air.	$0 - 10$	Rainy	Gauge	Gauge		different Points of Azimuth.							ಕ್ಷ	Pressure in lbs. on	
						Days.	read Daily.	read Monthly.	N.	N.E.	Е.	S.E.	S.	S.W.	W.	N.W.	Number nearly	the Square Foot.	<i>Ifean</i> Daily Horizontal Movement of Air in Miles. s
			85		.6.5	15	in. 1.49	in. $\overline{1}$ $\overline{5}8$	54	125	36	81	101	223	68	12			
$January \ldots \ldots$				555 555		13	0.54	0.60	34	192	52	62	135	111	4 ³	36	44	o•6q	297
February $\text{March} \dots \dots \dots$		78 83		554	7.4 7.5	11	2.05	2.15	130	262	28		48	115	52		7	0.88	35 ₉
April		6q		546		6	0.38	0.24	61.	34	84	19 51				87	3	0.42	310
$May \ldots \ldots \ldots$				539	4° 5.6	5		0.37	63	105	63	66	44 3 ₁	201	118	84	43 28	0'17	239
June.			7 ³ 68		6.4		0.47 o.39	0.23	96		2I	25		298	46	44 93		0.41	255
			532		6.6	4			60	97	67		14 35	193	177	61	$\overline{4}$ 6	0.26	243
$July \dots \dots \dots \dots$		70		525		10	2.01	1.93		115		42		217	141			0.18	223

MONTHLY MEANS of RESULTS for METEOROLOGICAL ELEMENTS.

 $7³$ August $\dots \dots$ 529 6.2 $\mathbf{1}\circ$ 2.02 2^{\degree} 00 164 185 $23|16|118$ $81 | 105$ $\bf 8$ 44 0.39 240 $9^{\rm 5}$ September....... 83 1.63 56 46 537 4.3 $\boldsymbol{8}$ $1'70$ 19 109 122 218 27 $\bf 28$ $o.36$ 229 $October \dots \dots$ 85 537 6.2 16 3.34 62 25 55 3.40 48 7° 264 156 64 \circ $\circ 65$ 327 November 90° 548 6.1 $9₁$ 1.50 1.30 $77\,$ 67 $4³$ 60 67 238 121 $\bf 28$ 19 0.34 240 December 65 86 559 $\mathbf{8}\cdot\mathbf{1}$ $3 \cdot 13$ 3.00 138 69 $\bf 2 \, 2$ 210 23 106 $7^{\rm 3}$ 3_I 29 0.31 242 Sum $\operatorname{\mathbf{Sum}}$ Sum $\operatorname{\mathbf{Sum}}$ Sum Sum Sum Sum Sum Sum $\operatorname{\mathbf{Sum}}$ $\operatorname{\mathbf{Sum}}$ Means 543 6.2 79 129 18.55 18.20 1571 691 $|575|615|$ 2302 0.41 267 944 1171 670 221

(I.)-Reading of a Thermometer whose bulb is sunk to the depth of 25'6 feet (24 French feet) below the surface of the soil, at Noon on every Day, except Sundays and Good Friday.

(II,)-Reading of a Thermometer whose bulb is sunk to the depth of 12 '8 feet (12 French feet) below the surface of the soil, at the same times.

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(II,)-.-Reading of a Thermometer whose bulb is sunk to the depth of 12 '8 feet (12 French feet) below the surface of the soil, at the same times-concluded.

(III.)-Reading of a Thermometer whose bulb is sunk to the depth of 6.4 feet (6 French feet) below the surface of the soil, at the same times,

At temperatures below 43° 5 the fluid of this thermometer descends below the scale; the readings from February 23 to March 5 were less than 43° 5. GREENWICH OBSERVATIONS, 1870.

(III.)-Reading of a Thermometer whose bulb is sunk to the depth of 6' 4 feet (6 French feet) below the surface of the soil, at the same times-concluded.

(IV,)-Reading ofa Thermometer whose bulb is sunk to the depth of 3'2 feet (3 French feet) below the surface of the soil, at the same times,

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(V,)-Reading of a Thermometer whose bulb is sunk to the depth of I inch below the surface of the soil, within the case which covers the tops of the deep-sunk Thermometers, at the same times,

(VI.)-Reading of a Thermometer within the case covering the deep-sunk Thermometers, whose bulb is placed on a level with their scales, at the same times,

(Ixxxiv) READINGS OF THERMOMETERS SUNK IN THE GROUND,

(VI.)-Reading of a Thermometer within the case covering the deep-sunk Thermometers, whose bulb is placed on a level with their scales, at the same times-concluded.

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(lxxxvi) CHANGES OF THE DIRECTION OF THE WIND,

ABSTRACT OF THE CHANGES OF THE DIRECTION OF THE WIND, AS DERIVED FROM OSLER'S ANEMOMETER.

- By *direct* motion, in the following statements, is meant that the change of the direction of the wind was in the order N., E., S., W., N., &c.; by *retrograde* is meant in the order N., W., S., E., N., &c.
- 186g. Dec. 31.12. The direction of the wind was S.

d h

1870. Jan. 31. 12. " "S.S.W., which implies a direct motion of $382\frac{1}{2}$.

On Jan. 15. 21. 30^m, 16^d. 22^h, 26^d. 22^h, 28^d. 8^h, 45^m, the trace was shifted to the next set of lines upwards, implying retrograde motion of 1440°.

Therefore the whole excess of retrogade motion in the month of January was $1057\frac{1}{5}$.

1870. Jan. 31. 12. The direction of the wind was $S.S.W.$

Feb. 28. 12. ", \qquad , $\$

On Feb. 2.22, the trace was shifted to the next set of lines downwards; on Feb. q^d . 21^b , 13^d , q^h . 15^m , the trace was shifted to the next set of lines upwards, implying direct motion of 360°, and retrograde motion of 720°.

Therefore the whole excess of retrograde motion in the month of February was $742\frac{1}{5}$ °

1870. Feb. 28.12 . The direction of the wind was S.

March 31. 12. \qquad ,, E., which implies a direct motion of 270° .

On March 2. 22, 4^d . 2^h , 45^m , 12^d . 22^h , 25^d , 9^h . 15^m , 31^d . 2^h , 45^m , the trace was shifted to the next set of lines upwards; on March 3^d, 3^h, ^Igd. 22h, the trace was shifted to the next set of lines downwards, implying retrograde motion of 1800°, and direct motion of 720°.

Therefore the whole excess of retrograde motion in the month of March was 810°.

1870. March 31. 12. The direction of the wind was E.

On April 18.2. 45^m , the trace was shifted to the second set of lines downwards; and on April 1^d. 22h, 6^d, 22h, 27^d. 20h. 45^m, to the next set of lines downwards; on April 1^d , 2^h , 4^5 m, 2^d . 8^h , 4^5 m, 7^d . 2^h , 4^5 m, 1^d . 2^h , 4^5 m, 1^d , 2^h , 4^5 m, 1^d , 2^h , 4^5 m, 1^d , 2^h , 4^h , 3^h , 3^h , to the next of lines upwards, implying direct motion of 1800°, and retrograde motion of 1800°.

Therefore the whole excess of retrograde motion in the month of April was 2021°.

- 1870. April 30. 12. The direction of the wind was $W.S.W.$
- May $31.12.$ " $\mathbf{W.S.W.}$, which implies no change.
- On May 4.21.30^m, 14^d. 23h. 45^m, 23^d. 0h. 10^m, 27^d. 23h. 45^m, 28d. 22h, the trace was shifted to the next set of lines downwards; on May 9^d. 9^h. 15^m, 15^d. 21^h. 15^m, 22^d. 9^h. 15^m, 23^d. 2^h. 45^m, 29^d. 23^h. 45^m, the trace was shifted to the next set of lines upwards, implying direct motion of 1800°, and retrograde motion of 1800°.

Therefore there was no change in the month of May.

1870. May 31.12. The direction of the wind was W.S.W.
June 30.12. \longrightarrow W , which implies a direct motion of 22¹/₂.

June 30. 12.
(On June 28. 8. 45^m, the trace was shifted to the second set of lines upwards; and on June 6^d. 22^h, 14^d· 22^h, 16^d· 2^h· 45^m, 16d²· 8^h· 30^m, to the next set of lines upwards; on June 3d. 22h, 8d. 8h. 45m, 14d. 8h. 45m, 16d. 20h. 45m, 20d. 8h. 45m, 21d. 2h. 45m, 22^d . 2^h , 45^m , 22^d . 8^h , 45^m , 23^d . 22^h , the trace was shifted to the next set of lines downwards, implying retrograde motion of 2160°, and direct motion of 3240°.

Therefore the whole excess of direct motion in the month of June was $1102\frac{1}{2}$.

1870' June 3o. 12. The direction of the wind was W.

July 31.12. \ldots , \ldots , $N.E., which implies a direct motion of 135°.$

On July 7. 2. 45^m , 9^d . 2^h , 45^m , 11^d . 20^h , 45^m , 21^d . 22^h , 24^d . 8^h , 0^m , 25^d . 22^h , 29^d . 8^h , 45^m , the trace was shifted to the next set of lines downwards; on July 21^d. 23^h. 45^m, the trace was shifted to the second set of lines upwards; and on July 7^d. 20^h. 45^m, 8^d . oh. 20^m, 8^d . 9^h . 30^m , 8^d . 22^h , 24^d . 22^h , 27^d . 8^h . 45^m , to the next set of lines upwards, implying direct motion of 2520^o , and retrograde motion of 2880°.

Therefore the whole excess of retrograde motion in the month of July was *225°.*

April 30. 12. ", which implies a retrograde motion of $202\frac{1}{2}$.

1870. July 31. 12. The direction of the wind was N.E.

Aug. 31. 12. " , $S.S.W.,$ which implies a retrograde motion of 202 $\frac{1}{2}$.

On Aug. 3. 20. 45^{m} , 6d. 2^{h} , 45^{m} , 7^{d} , 22^{h} , 8^{d} , 8^{h} , 45^{m} , the trace was shifted to the next set of lines upwards; on Aug. 14d, 9h. 45m, 17^d , 21^h , 21^h , 30^m , 25^d , 45^m , the trace was shifted to the next set of lines downwards, implying retrograde motion of 1440°, and direct motion of 1440°.

Therefore the whole excess of retrograde motion in the month of August was $202\frac{1}{5}$.

1870. Aug. 31. 12. The direction of the wind was S.S.W.

Sept. 30. 12. \ldots , , E.N.E., which implies a direct motion of 225°.

On Sept. 23. 8. 45^{m} , 29^d. 0^{h} , the trace was shifted to the second set of lines upwards; and on Sept. 18^d. 22^h, 19^d. 22^h, 20^d. 0^{h} , 2^{d} . 2^{h} . 45^{m} , 2^{d} . 2^{h} , to the next set of lines upwards; on Sept. 16^d. 21^{h} , 17^{d} . 23^{h} . 30^{m} , 20^{d} . 8^{h} . 30^{m} , 29^{d} . 8^{h} . 45^{m} , to the next set of lines downwards, implying retrograde motion of 3240°, and direct motion of 1440°.

Therefore the whole excess of retrograde motion in the month of September was 1575°.

d h

1870. Sept. 30. 12. The direction of the wind was E.N.E.

Oct. 31.12. ", , N., which implies a retrograde motion of $67\frac{1}{2}$.

On Oct. 11. 22, the trace was shifted to the next set of lines upwards; on Oct. 14^d. q^h . 15^m, 14^d. 20^h. 45^m, the trace was shifted to the next set of lines downwards, implying retrograde motion of 360°, and direct motion of *7zoo.*

Therefore the whole excess of direct motion in the month of October was $2q_2\frac{10}{2}$.

d h 1870' Oct. 31. 12. The direction of the wind was N.

Nov. 30.12. \qquad ,, \qquad , \qquad , \qquad , \qquad S.E., which implies a direct motion of 135°.

On Nov. 2. 22, 5^d, 22^h, 9^d, 0^h, the trace was shifted to the next set of lines downwards; on Nov. 5^d, 23^h, 45^m, 7^d, 21^h, 27^d, 20^h, 30^m, the trace was shifted to the next set of lines upwards, implying direct motion of 1080°, and retrograde motion of 1080°.

Therefore the whole excess of direct motion in the month of November was 135°.

d h 1870. Nov. 30. 12. The direction of the wind was S.E.

Dec. 31.12. ", $\qquad \qquad$, $\qquad E.N.E., which implies a retrograde motion of $67\frac{1}{2}^\circ$.$

On Dec. 9. 22, 15^d. 20^h. 45^m, the trace was shifted to the next set of lines upwards; on Dec. 2^d. 0^h. 20^m, 12^d. 22^h, the trace was shifted to the next set of lines downwards, implying retrograde motion of 720°, and direct motion of 720°.

Therefore the whole excess of retrograde motion in the month of December was $67\frac{1}{6}$.

The whole excess of retrograde motion to the end of the year was $3352\frac{1}{2}$.

The revolution-counter which is attached to the vertical spindle of the vane, whose readings increase with change of direction of the wind in the order N., E., S., W., &c., or in *direct* motion, and decrease with change of direction in the order N., W., S., E., &c., or in *retrograde* motion, gave the following readings :-

(lxxxviii) AMOUNT OF RAIN COLLECTED IN EACH MONTH.

AMOUNT OF RAIN COLLECTED IN EACH MONTH OF THE YEAR 1870.

The heights of the receiving surfaces are as follows:

ROYAL OBSERVATORY, GREENWICH.

OBSERVATIONS

OF

LUMINOUS METEORS.

1870.

GREENWICH OBSERVATIONS, 1870. **And All According to the CONTEX CONTROL**

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(xc) OBSERVATIONS OF LUMINOUS METEORS,

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No. for Reference. Path of Meteor through the Stars. IFrom a point nearly midway between α Cephei and α Lyrre at an inclination of 45°.
 EXECUTE: From α Orionis to f Tauri. From q Orionis to f Tauri. .3 From the Pleiades disappeared about 5° below Jupiter. 4 Fell almost perpendicularly in the S. No stars near. Moon about 25° above and somewhat to the left of point of appearance of From a point near l Orionis, passed close to κ Orionis, and burst midway between α and From a point near l Orionis, passed close to κ Orionis, and burst midway between α and β Canis Majoris. 6 From direction of *p* Lyncis disappeared close to Procyon. 7 From direction of β Cephei passed close below Cassiopeia, and disappeared a little below β Pegasi.
8 From α Leonis disappeared about γ° below Procyon. From α Leonis disappeared about 7° below Procyon. 9 From direction of Polaris to , Auriga. 10 From a point about 5° to the left of β Geminorum passed nearly across β Canis Minoris. **II From the direction of Polaris passed midway between** α **Aurigæ and** α **Geminorum.
I2 From a point near L Camelopardali disappeared close to** β **Aurigæ.** From a point near L Camelopardali disappeared close to β Aurigæ. 13 | Passed between ϵ and δ Virginis from direction of ζ Bootis. 14 From η Virginis disappeared a little below α Virginis.
15 Directed from γ Coronae Borealis disappeared near ξ 1 15 Directed from γ Coronæ Borealis disappeared near ξ Bootis.
16 In West moving horizontally from N. to S. at altitude 20°. 16 In West moving horizontally from N. to S. at altitude 20°. Below, and to the left of α Leonis.
17 From direction of γ Draconis towards α Coronæ Borealis.
18 Passed close to β Cephei from direction of θ C 17 | From direction of γ Draconis towards α Corone Borealis. 18 Passed close to β Cephei from direction of θ Cephei. 19 From a point near η Bootis fell slowly to τ Virginis. 20 From α Bootis to a point midway between α and κ Virginis. 21 From direction of Polaris passed towards a point 10° W. of β Aurigæ. 22 From direction of Polaris towards α Auriga. 23 From direction of Polaris passed about 1° to the left of θ Geminorum. 24 | From direction of κ Ursæ Majoris fell towards α Geminorum. 24 From direction of κ Ursæ Majoris fell towards α Geminorum.
25 From direction of Polaris fell perpendicularly towards N. horizon. 26 From the direction of α Coronæ, fell at right angles to a line joining α Lyræ and Arcturus. 27 From a point between α and β Aurigæ fell vertically towards horizon. 28 From direction of β Ophiuchi burst near α Cephei. 29 From ζ Cephei to δ Cygni. 30 From direction of *a* Cassiopeixe to a little below β Ursx Majoris. 31 Fell at inclination of 7° from perpendicular, from near γ Andromedæ. 32 Fell from δ Andromedæ to the horizon at an angle of 45° to the left.
33 From a point a few degrees below α Cephei passed close to α Lyræ. 33 From a point a few degrees below α Cephei passed close to α Lyræ. 34 From a point about 20° below *a* Lyrre fell vertically towards *a* Aquarii. 34 From a point about 20° below *a* Lyræ fell vertically towa
35 From a point a few degrees below *a* Pegasi fell vertically.
36 From a point about 15° North of *a* Persei shot in the direct 36 From a point about 15⁶ North of a Persei shot in the direction of δ Ursæ Majoris. 37 From a point about 12° below *a* Pegasi fell at an angle of 30° to the left towards the horizon. 37 From a point about 12° below α Pegasi fell at an angle 38 From β Cephei disappeared between o and κ Honorum. 39 From a point close before α Pegasi to β Cephei. 40 From a point about 5° N. of α Persei disappeared about 10° in front of γ Cephei. 40 From a point about 5° N. of α Persei disappeared about 10° in 1
41 From ζ Cassiopeix to a point about 5° to right of α Cassiopeix.
43 From θ Cassiopeix to α Honorum. 42 From a point about 5° below a Persei shot upwards at an inclination of 85° (left of vertical). 43 From θ Cassiopeixe to *k* Honorum.
44 From a point about 5° below δ Cassiopeixe to ζ Draconis.
46 From *k* Persei shot to a point between From a point about 5° below δ Cassiopeix moved nearly parallel to δ and α Cassiopeix. From β Cassiopeixe to ζ Draconis. 46 From κ Persei shot to a point between γ and β Andromedæ.
47 From a point a little above γ Andromedæ moved in the dire
48 From a point a few degrees below β Cephei passed midway From a point a little above γ Andromedæ moved in the direction of α Andromedæ. From a point a few degrees below β Cephei passed midway between ϵ and δ Cassiopeiæ. 49 From κ Persei across α Ursæ Majoris.
50 From direction of β Cassiopeiæ toward From direction of β Cassiopeixe towards Polaris. 51 Center of path below and opposite β Ursæ Majoris; moving from direction of α Draconis. 52 From a point about 10° above a Persei fell towards horizon at angle of 25° to left. $\overline{53}$ | From direction of *a* Pegasi passed across *a* Andromedæ towards $\tilde{\beta}$ Andromedæ. 54 From a point a few degrees to left of γ Persei towards α Aurigæ. 55 Passed across ϵ Cygni from direction of β Cygni.

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OBSERVATIONS OF LUMINOUS METEORS,

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(xciv) OBSERVATIONS OF LUMINOUS METEORS,

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OBSERVATIONS OF LUMINOUS METEORS,

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OBSERVATIONS OF LUMINOUS METEORS,

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 $\label{eq:2} \frac{1}{2} \sum_{i=1}^N \frac{1}{2} \sum_{j=1}^N \frac{1}{$ $\label{eq:2.1} \frac{1}{2} \int_{0}^{2\pi} \frac{1}{\sqrt{2}} \, \frac$ $\label{eq:2} \frac{1}{2} \int_{\mathbb{R}^3} \frac{d^2\mathbf{r}}{|\mathbf{r}|^2} \, \mathrm{d}\mathbf{r} \, \mathrm$ $\mathcal{L}^{(1)}$ and $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{\mathbb{R}^3}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^2\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}\frac{1}{\sqrt{2}}$ $\frac{1}{\sqrt{2}}$ $\label{eq:2.1} \begin{split} \frac{d\mathbf{x}}{d\mathbf{x}}&=\frac{d\mathbf{x}}{d\mathbf{x}}\left(\frac{d\mathbf{x}}{d\mathbf{x}}\right)\\ &=\frac{d\mathbf{x}}{d\mathbf{x}}\left(\frac{d\mathbf{x}}{d\mathbf{x}}\right)\\ &=\frac{d\mathbf{x}}{d\mathbf{x}}\left(\frac{d\mathbf{x}}{d\mathbf{x}}\right)\\ &=-\frac{d\mathbf{x}}{d\mathbf{x}}\left(\frac{d\mathbf{x}}{d\mathbf{x}}\right)\\ &=-\frac{d\mathbf{x}}{d\mathbf{x}}\left(\frac{d\mathbf{x}}{d\math$ $\label{eq:1} \begin{aligned} \mathbf{E}^{(1)}_{\text{max}} &= \frac{1}{2} \sum_{i=1}^{N} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(2)}_{\text{max}} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(2)}_{\text{max}} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(2)}_{\text{max}} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(2)}_{\text{max}} \mathbf{E}^{(1)}_{\text{max}} \mathbf{E}^{(2)}_{\text{$ $\label{eq:2.1} \begin{split} \mathcal{L}_{\text{max}}(\mathbf{r}) = \frac{1}{2} \sum_{i=1}^{N} \mathcal{L}_{\text{max}}(\mathbf{r}) \mathcal{L}_{\text{max}}(\mathbf{r}) \\ \mathcal{L}_{\text{max}}(\mathbf{r}) = \frac{1}{2} \sum_{i=1}^{N} \mathcal{L}_{\text{max}}(\mathbf{r}) \mathcal{L}_{\text{max}}(\mathbf{r}) \end{split}$ $\label{eq:2} \frac{1}{\sqrt{2}}\int_{0}^{\pi}\frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^{2}d\mu_{\rm{eff}}$ $\mathcal{L}(\mathcal{A})$ and $\mathcal{L}(\mathcal{A})$ $\mathcal{A}^{\mathcal{A}}$

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 $\label{eq:2.1} \frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\right)\frac{1}{\sqrt{2\pi}}\right)\frac{d\omega}{\omega}d\omega.$ $\mathcal{L}^{\text{max}}_{\text{max}}$ and $\mathcal{L}^{\text{max}}_{\text{max}}$ $\label{eq:2.1} \frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^{2} \left(\frac{1}{\sqrt{2}}\right)^{2} \left(\$ $\label{eq:2.1} \frac{1}{\sqrt{2}}\left(\frac{1}{\sqrt{2}}\right)^{2} \left(\frac{1}{\sqrt{2}}\right)^{2} \left(\$